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# DEVELOPING AN ERROR MODEL FOR IONOSPHERIC PHASE DISTORTIONS IN L-BAND SAR AND INSAR DATA

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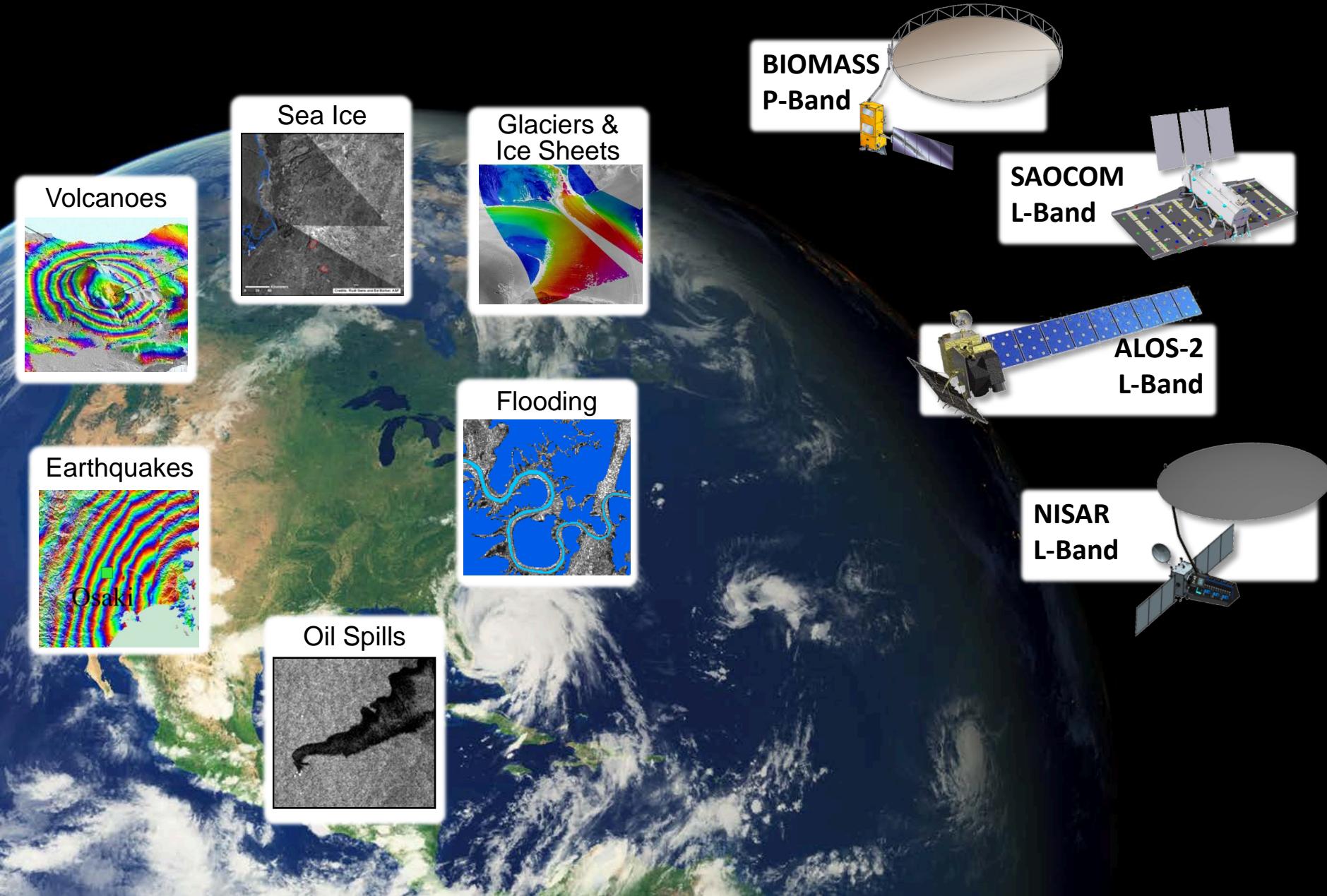
**UAF** ALASKA SATELLITE FACILITY  
*Making remote-sensing data accessible since 1991*



**JPL**

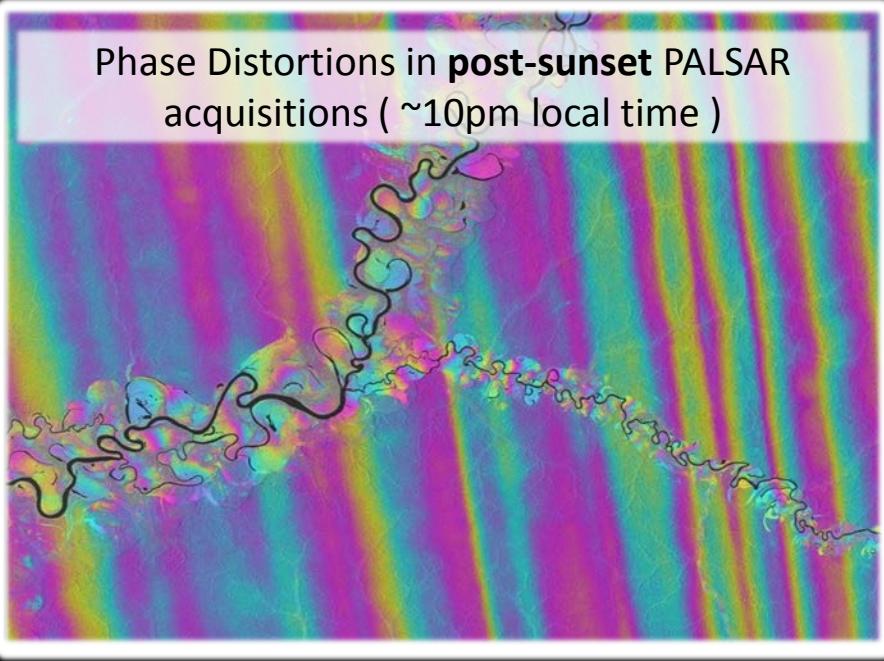


# Geophysical Applications Favor Low-Frequency SAR Systems

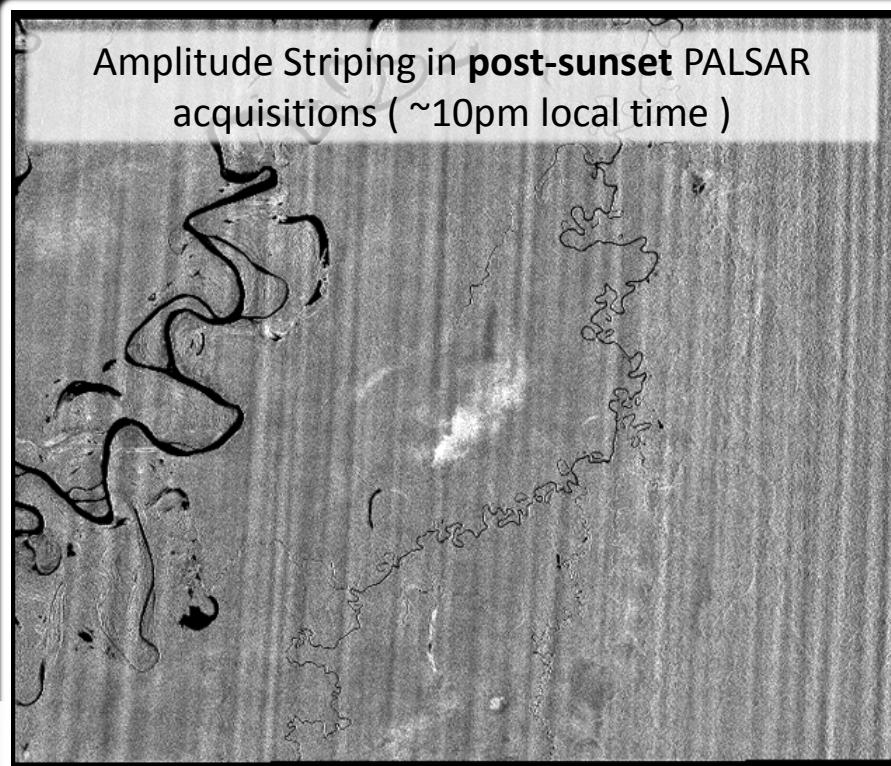


# Ionospheric Distortions in Low-Frequency (L-band) SAR

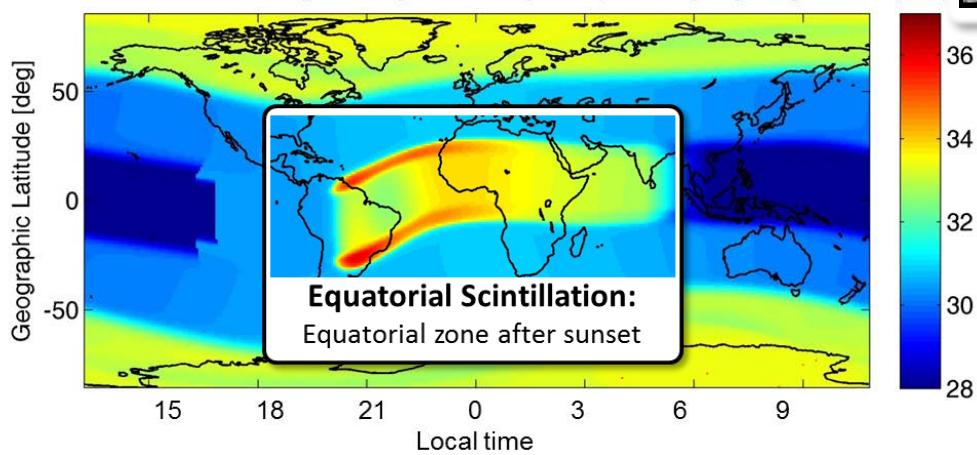
Phase Distortions in **post-sunset** PALSAR acquisitions (~10pm local time)



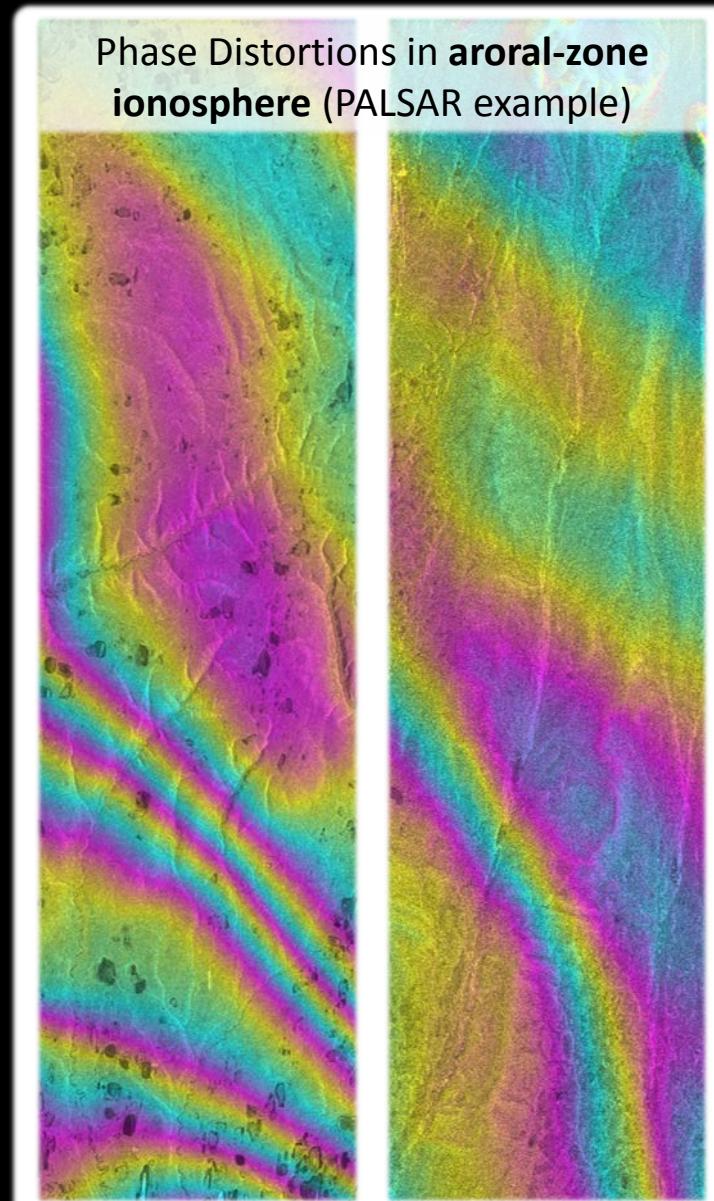
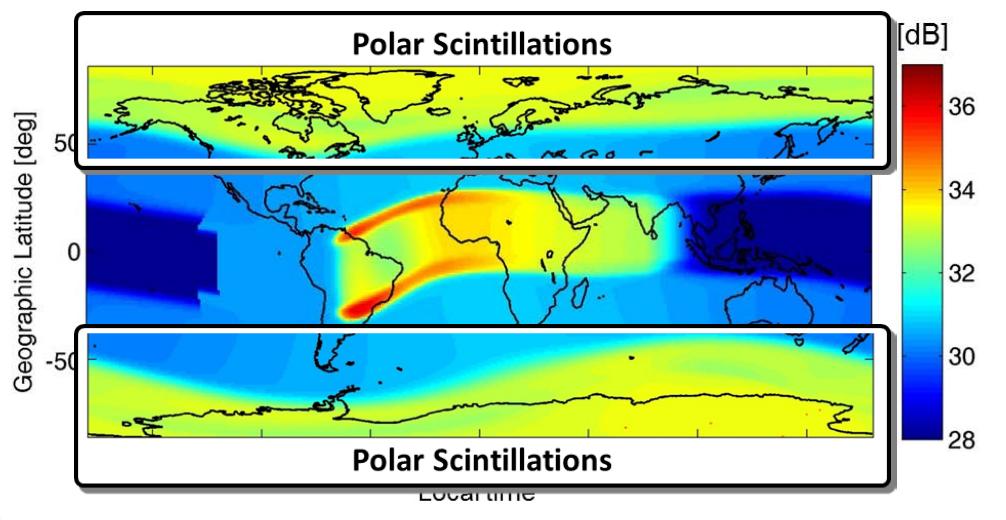
Amplitude Striping in **post-sunset** PALSAR acquisitions (~10pm local time)



L-band: Height Integrated Irregularity Strength [CkL]



# Ionospheric Distortions in Low-Frequency (L-band) SAR

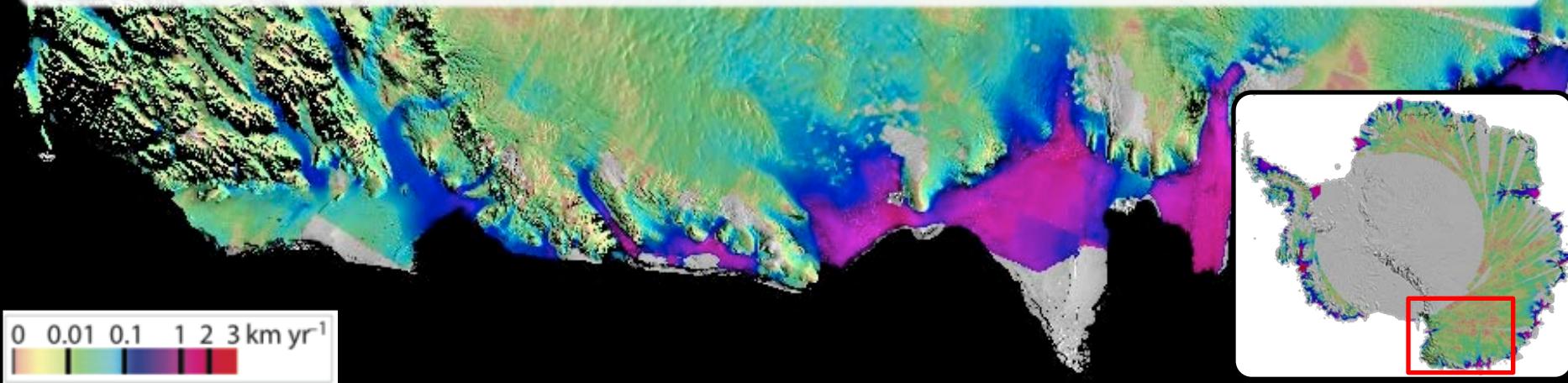


# Ionospheric Effects in L-band SAR-based Ice Velocity Measurements

Goal of this work: Develop a global ionospheric error model for

*Predicting ionospheric phase noise globally and on  
image-by-image basis*

- Essential for geodetic applications of SAR
- Useful information for mission design

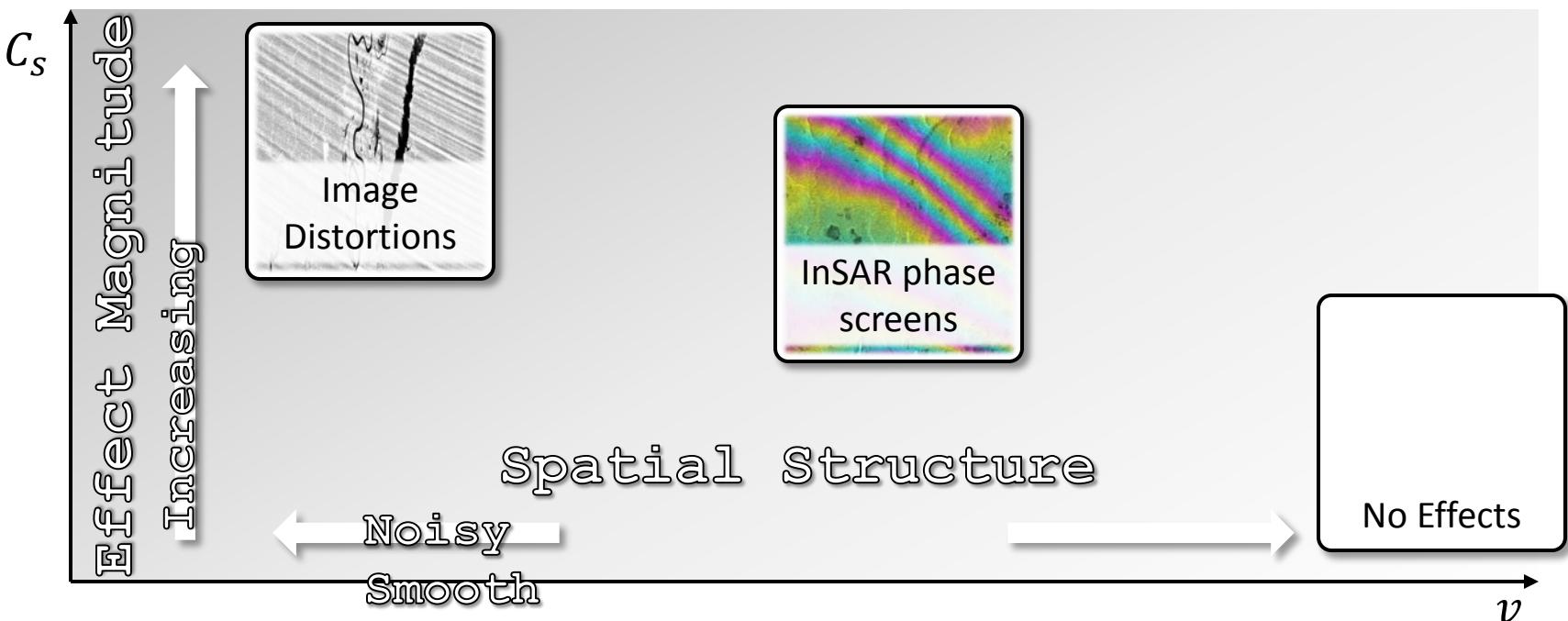


# Geophysical Origin of Observed Effects

- **Cause:** Ionospheric plasma irregularities due to scintillation
- **Effect:** Scintillation can be modeled using power law considerations

$$Q(k) = C_s \cdot (k_0^2 + k^2)^{-(\nu + \frac{1}{2})}$$

$k$ : spatial wavenumber;  $C_s$ : turbulence strength;  $\nu$ : spectral index;  $k_0$ : outer scale



# Geophysical Origin of Observed Effects

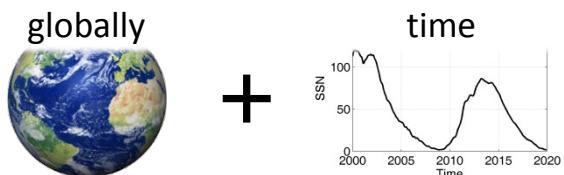
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## • Two Step Modeling Concept:

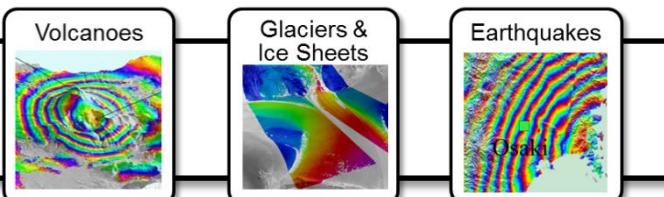
### 1. Predict ionospheric structure & calculate SAR phase power spectra



**Full characterization of ionospheric properties**  
(1) for each data take and (2) across mission lifetime

### 2. Propagate to phase variance & surface deformation errors

Determine effects on various earth science disciplines:



Volcanoes

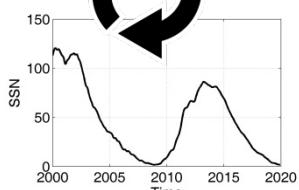
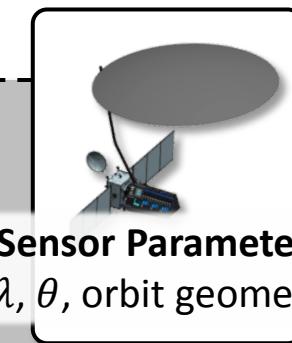
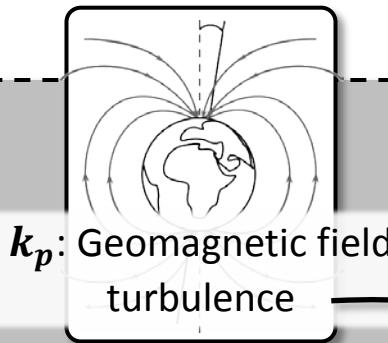
Glaciers &  
Ice Sheets

Earthquakes

# Deriving Ionospheric Phase Power Spectra

## Approach 1: Using Ionospheric Model WBMOD

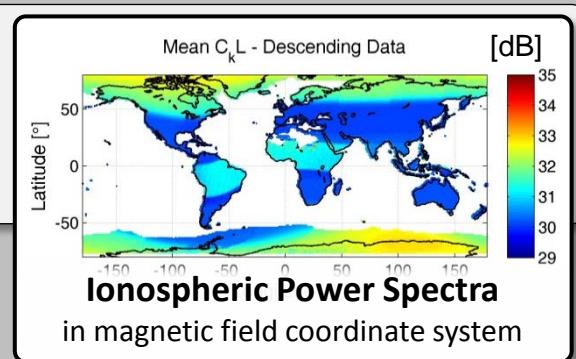
- Concept:



### WBMOD (Wide Beam MODel)

Climatological ionospheric model based on observations and models

*Rino & Ishimaru, Radio Science 27, 1992*



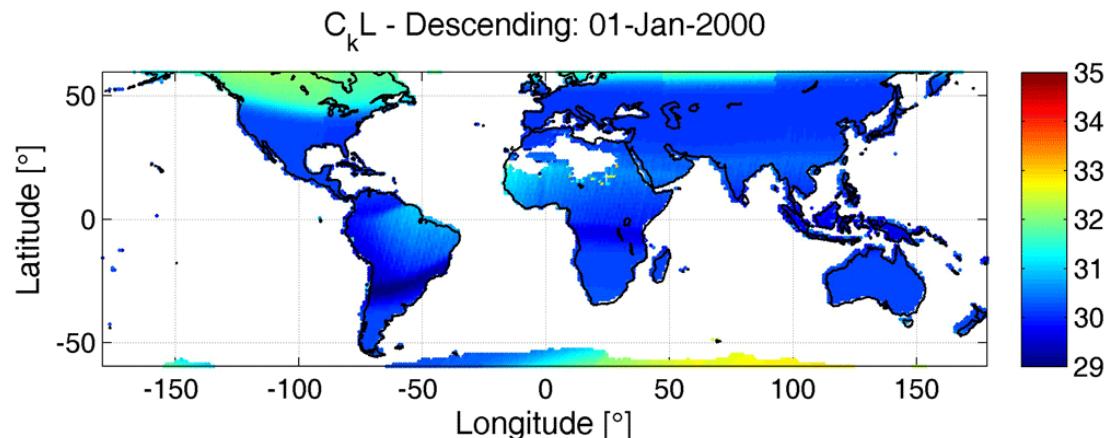
# Deriving Ionospheric Phase Power Spectra

## Approach 1: Using Ionospheric Model WBMOD

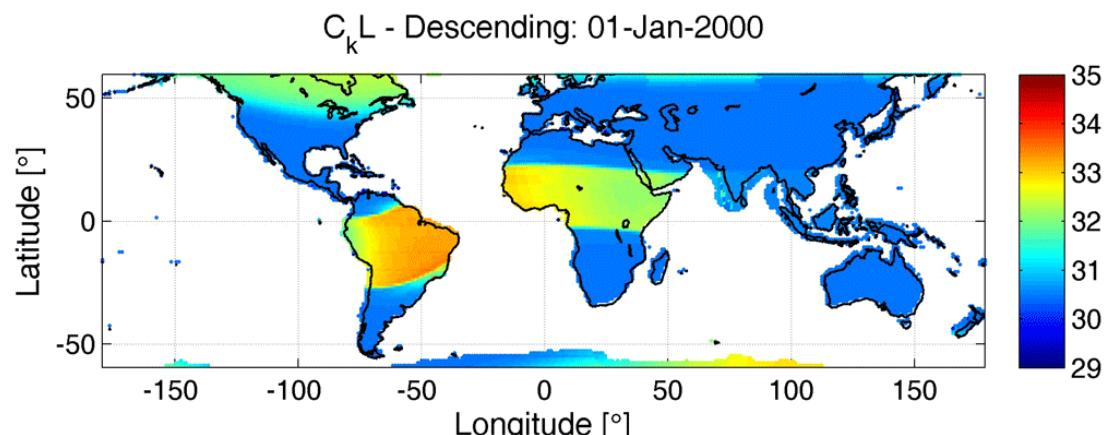
- **Mission-spanning simulations:**

- Frame-by-frame acquisition simulation every 5<sup>th</sup> cycle using satellite orbit information.
- Retrieval of relevant ionospheric parameters per frame and Averaging per  $1 \times 1^\circ$  grid cell

**NISAR CASE - DESCENDING**  
**acquisition time: 6pm**



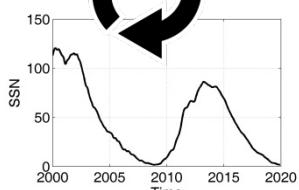
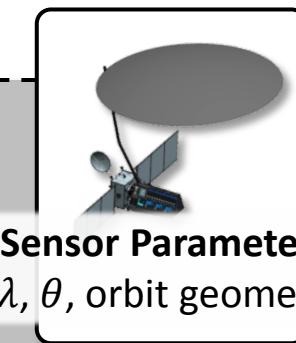
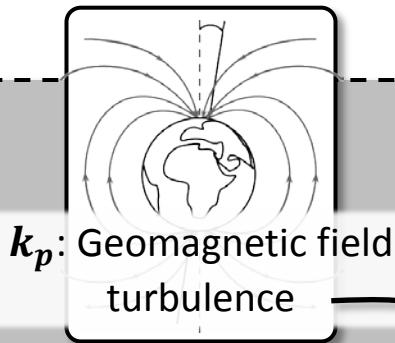
**ALOS-2 CASE - DESCENDING**  
**acquisition time: 12am**



# Deriving Ionospheric Phase Power Spectra

## Approach 1: Using Ionospheric Model WBMOD

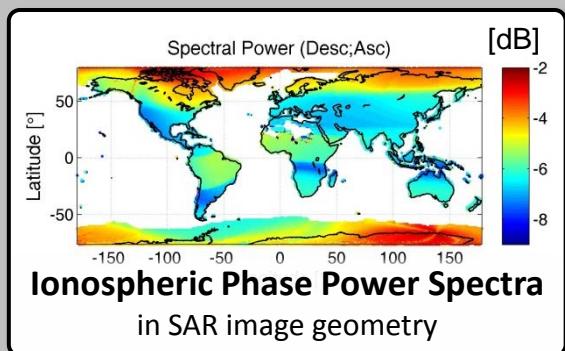
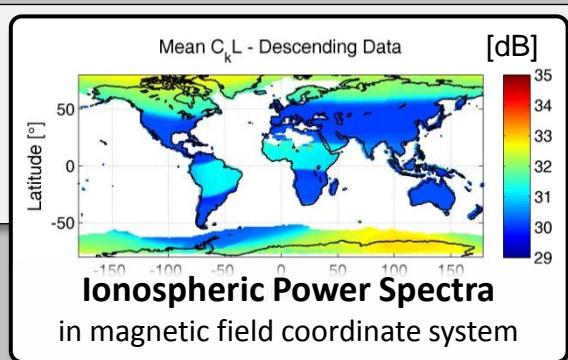
- Concept:



### WBMOD (Wide Beam MOdel)

Climatological ionospheric model based on observations and models

Rino & Ishimaru, Radio Science 27, 1992



### Projection into SAR Geometry

Meyer et al., 2014;  
Proceedings of EUSAR

# Deriving Ionospheric Phase Power Spectra

## Approach 1: Using Ionospheric Model WBMOD

- Transformation of ionospheric parameters into SAR phase power spectrum:

$$P_\phi(k, f) = \frac{r_e^2 \left(\frac{c}{f}\right)^2 \cdot \sec^2(\vartheta) \cdot a \cdot b \cdot C_k L \left(\frac{2\pi}{1000}\right)^{\nu+1}}{(k_0^2 + A k_x^2 + B k_x k_y + C k_y^2)^{(\nu+1/2)}}$$

The diagram illustrates the components of the equation:

- SAR parameters (sensor file)**:  $r_e^2 \left(\frac{c}{f}\right)^2 \cdot \sec^2(\vartheta) \cdot a \cdot b$
- Anisotropy parameters (from WBMOD)**:  $C_k L \left(\frac{2\pi}{1000}\right)^{\nu+1}$
- Ionospheric turbulence parameters (from WBMOD)**:  $\left(\nu+1\right)$
- Transformation factors (WBMOD)**:  $A k_x^2 + B k_x k_y + C k_y^2$
- geomagnetic east & west direction transverse wavenumbers (from WBMOD)**:  $k_0^2$

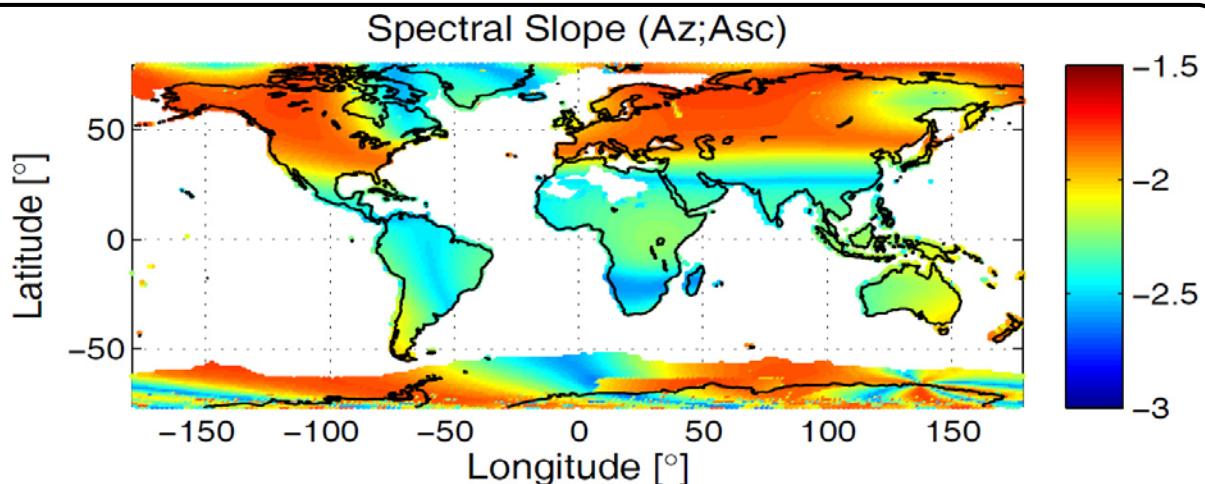
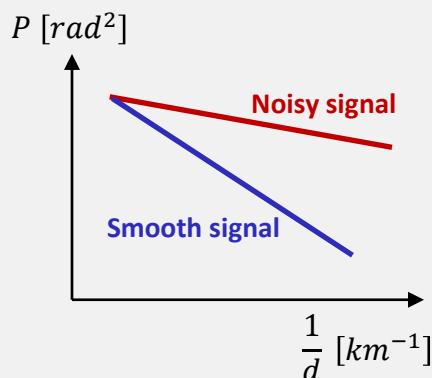
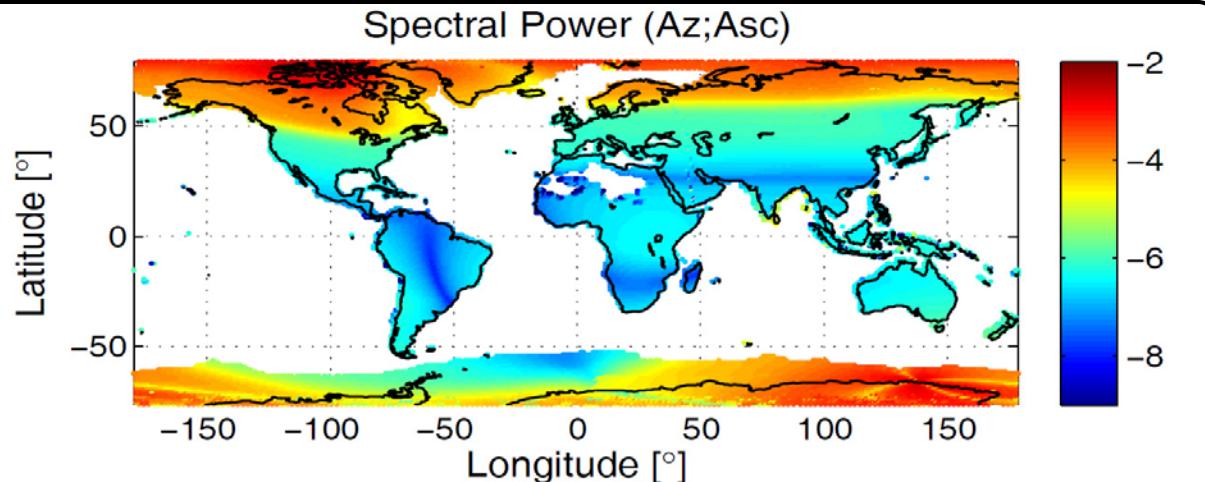
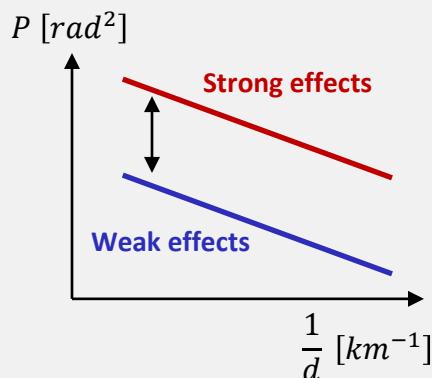
- Calculate average phase power spectrum parameters for mission lifetime

# Deriving Ionospheric Phase Power Spectra

## Approach 1: Using Ionospheric Model WBMOD

- Results: Ionospheric Phase Power Spectra in NISAR Data (Azimuth Direction)

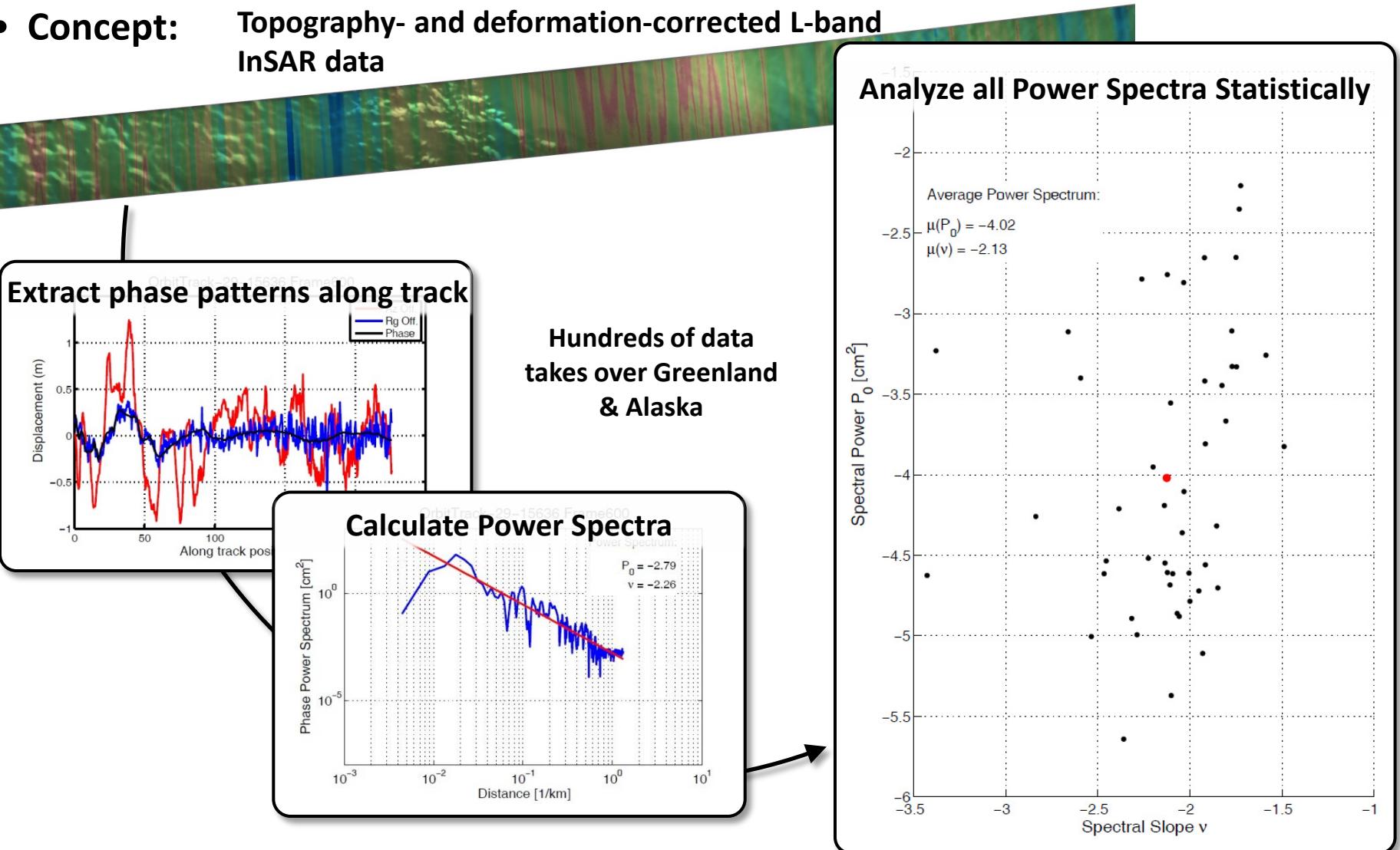
2.1 million model runs



# Deriving Ionospheric Phase Power Spectra

## Approach 2: Analysis of Available L-band SAR Archives

- Concept: Topography- and deformation-corrected L-band InSAR data



# Calculating Ionosphere-Induced Errors in Deformation Estimates

## Single Interferogram Analysis – The NISAR Case

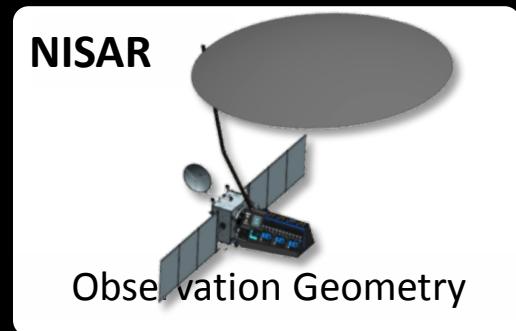
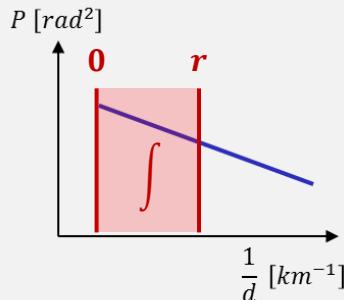
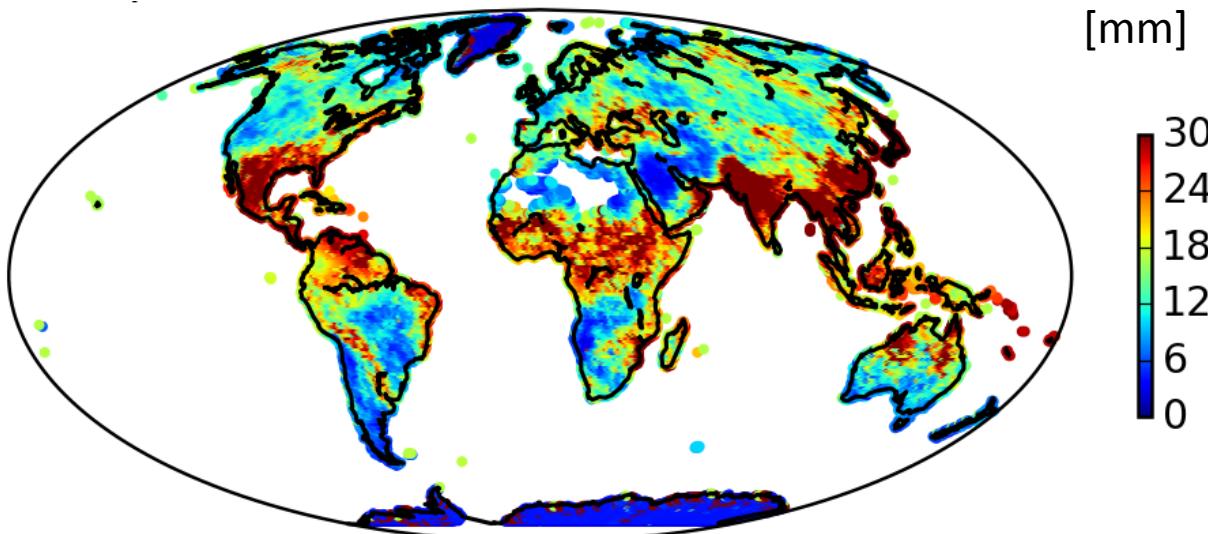


Image-by-Image Error Model ( $r = 50\text{km}$ )

Troposphere-only @ 50km scale

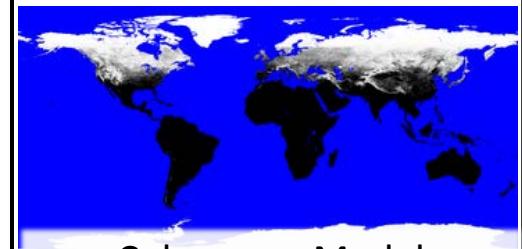


Typical Summer-time conditions

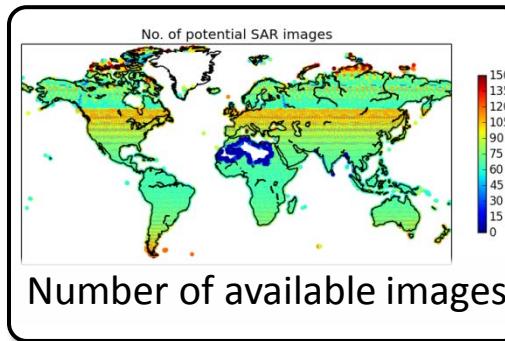
Tropospheric errors from MODIS analysis (Yun et al., AGU abstract G42A-02, 2014)

# Calculating Ionosphere-Induced Errors in Deformation Estimates

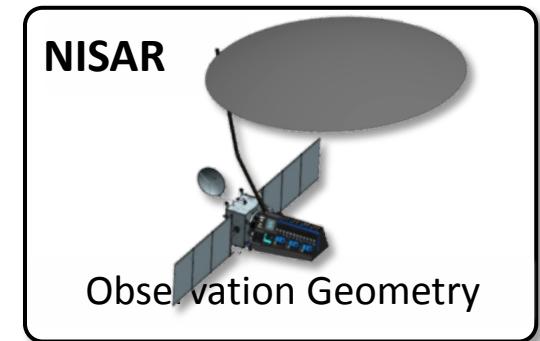
## Stack Processing – The NISAR Case



Coherence Model



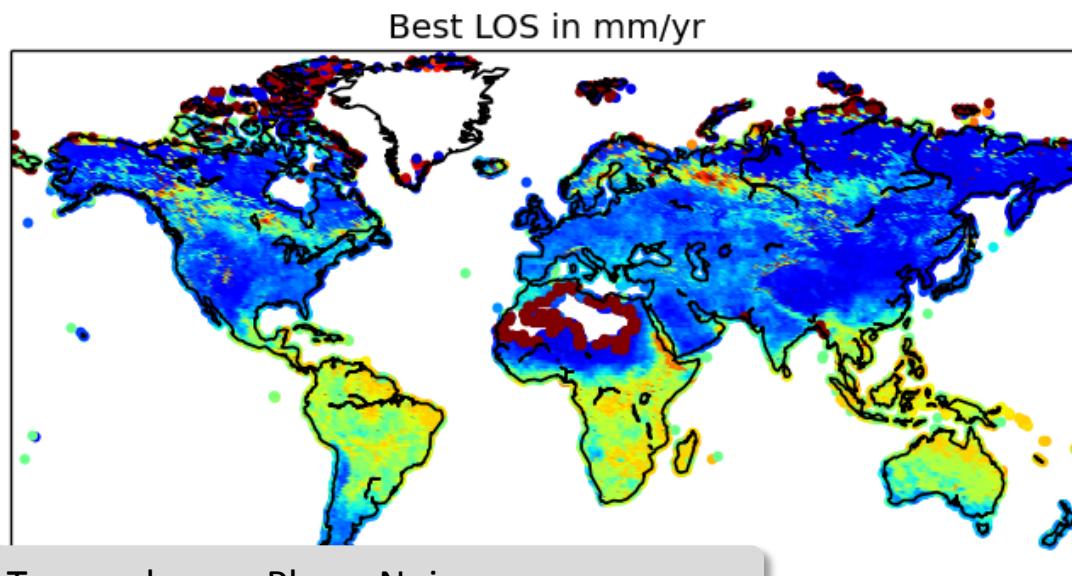
Number of available images



Observation Geometry

SBAS InSAR  
Model

Stack-based  
Deformation Error



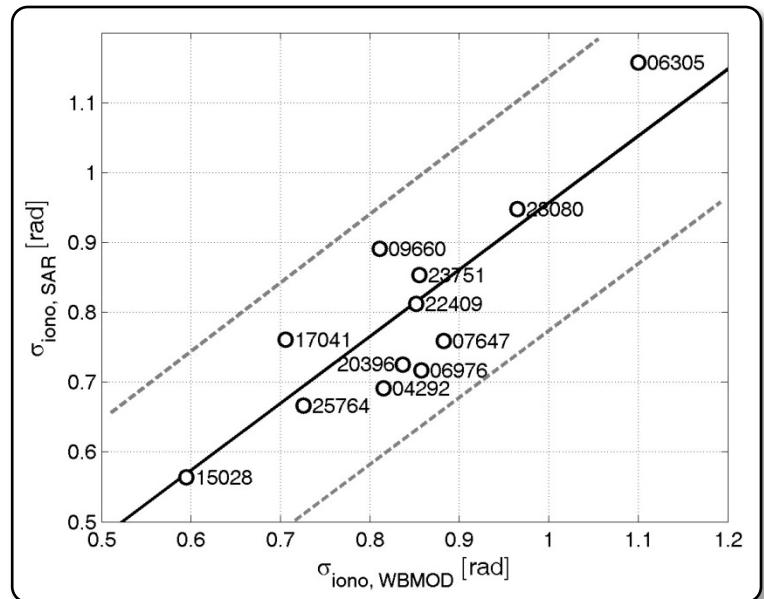
Best LOS in mm/yr

Troposphere + Phase Noise

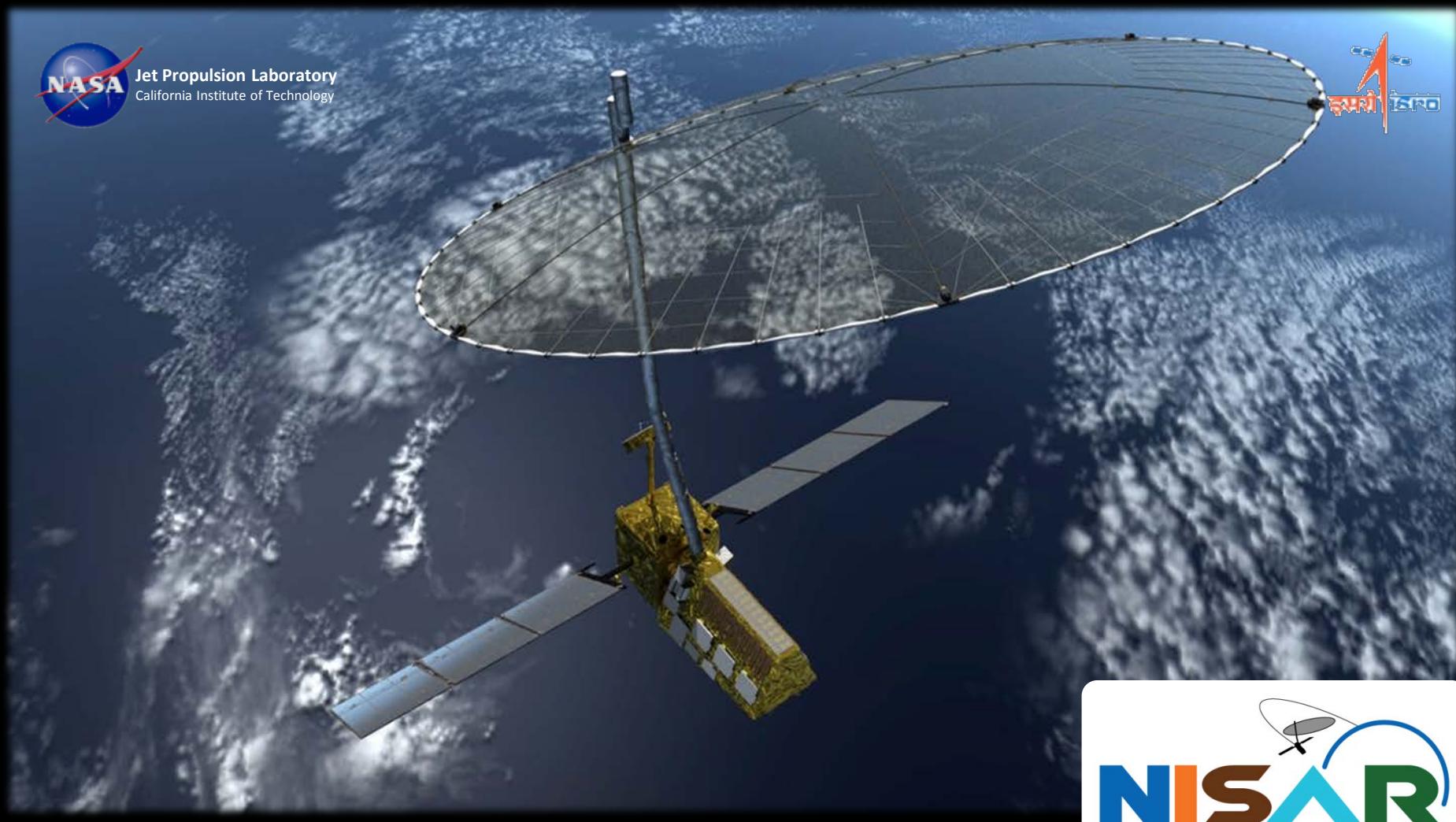
Output of JPL NISAR Performance  
Tool (Agram & Simons, JGR,  
accepted for publication)

# Conclusions

- Ionospheric effects are an issue in low-frequency SAR data
- A global ionospheric error model was developed that can be used in
  - Sensor and mission design
  - Image-by-image error analysis
- First model validation results indicate validity of the model
- Model was applied to NISAR mission; largest ionospheric influence in polar regions



# Thank you for your attention!



Funding provided by: NASA ROSES Grant No. NNX12AO29G  
Contact: [fjmeyer@alaska.edu](mailto:fjmeyer@alaska.edu) URL: [insar.alaska.edu](http://insar.alaska.edu)

