



TROPOMI

# TROPOMI SWIR on ground calibration

All results by Paul Tol, Richard van Hees, Matthijs Krijger,  
support Ralph Snel & OCAL Science Support team (KNMI)  
& Industrial Team (aDSnl, CSL, VSL)

Presented by Matthijs Krijger

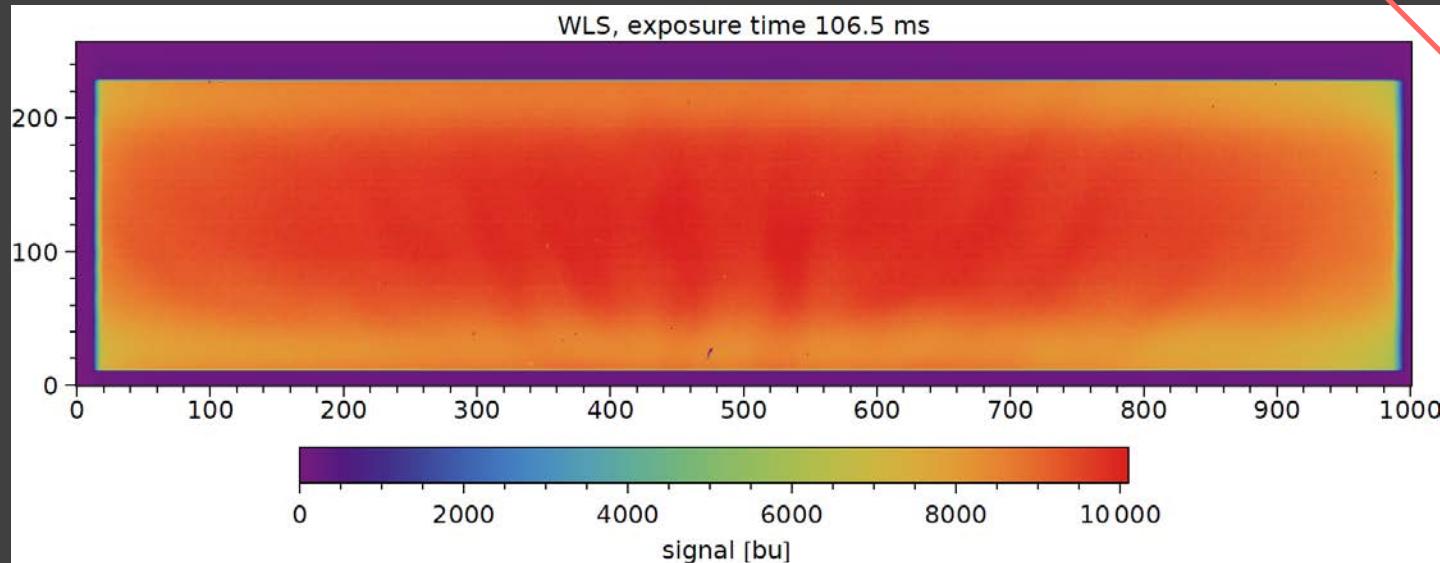


# TROPOMI SWIR requirements

For Lv2 SWIR retrievals (CO and CH<sub>4</sub>) the

- Instrument Spectral Response Function (ISRF) is required at high subpixel (0.02px) resolution
- Radiometric offset (straylight) must be reduced to less than 0.1% signal

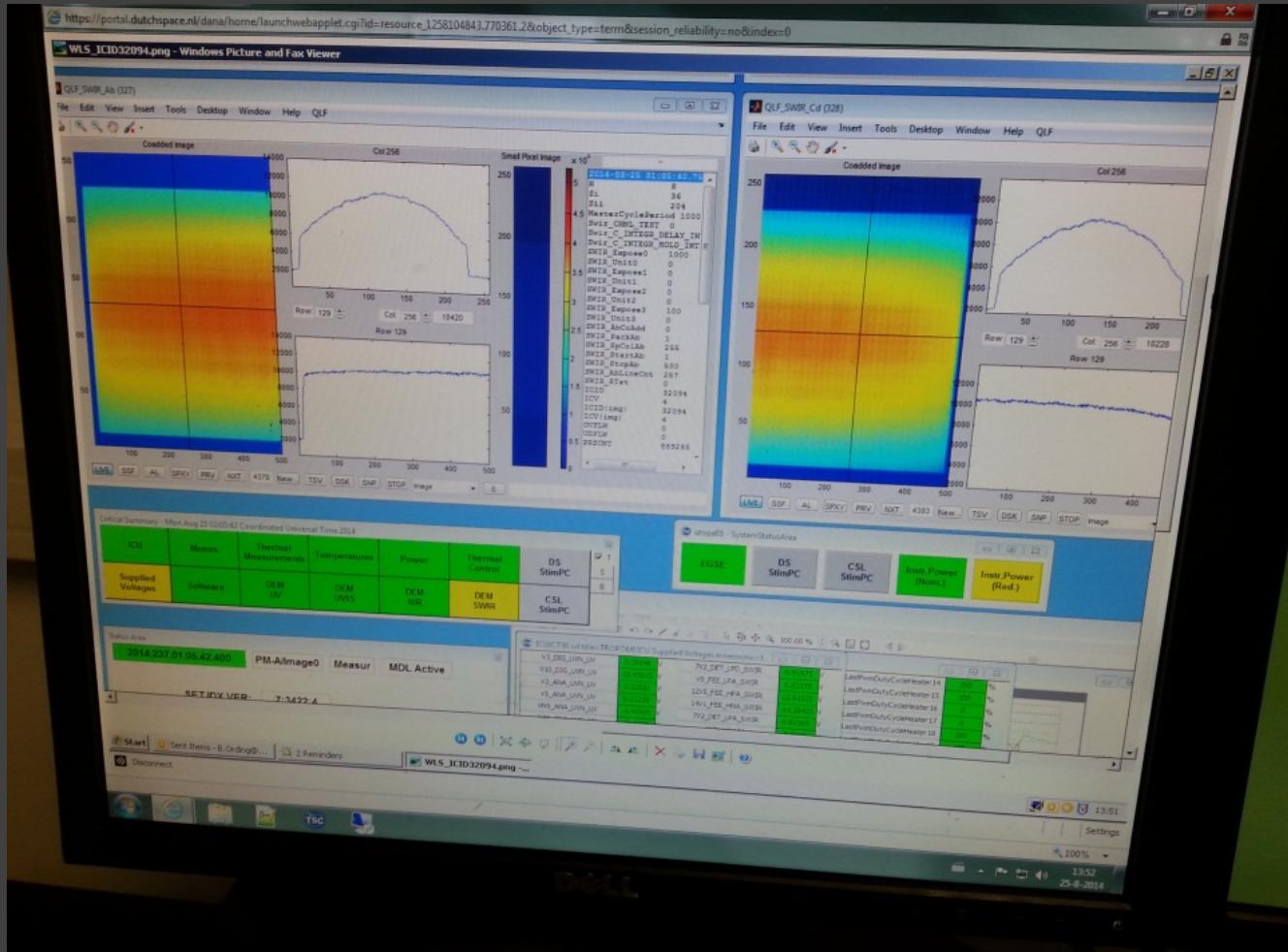
109 degrees  
[0.5 degree/px]  
viewing angles



# TROPOMI in 6.5m vacuum tank Centre Spatial Liege

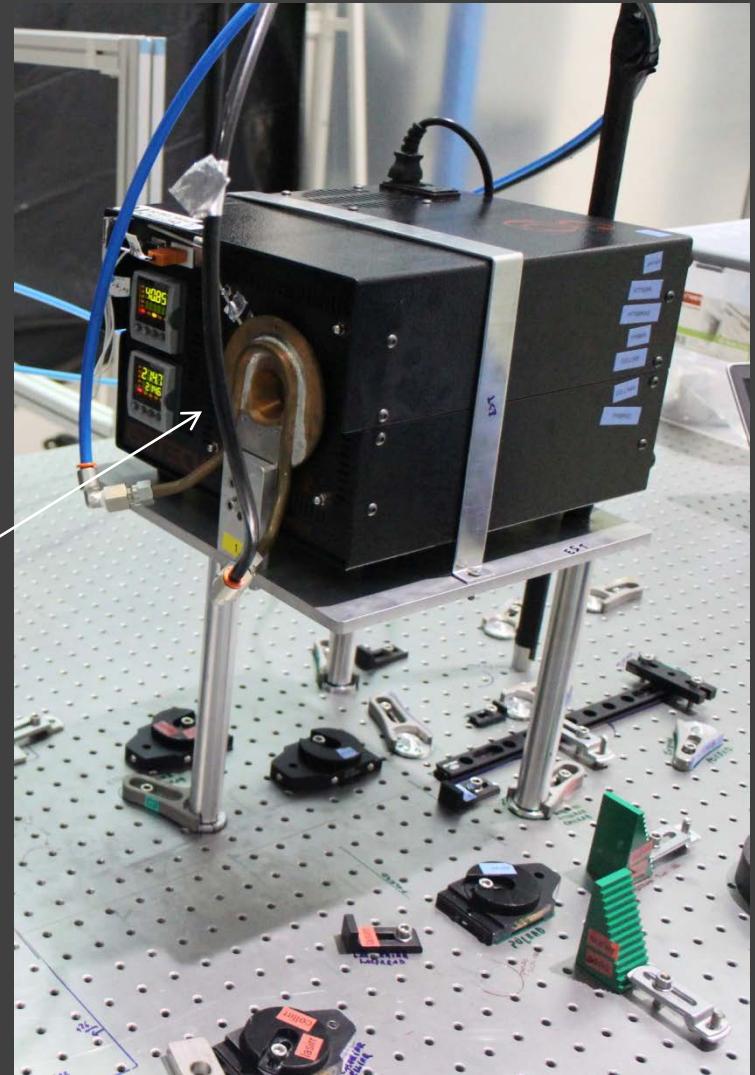


# First SWIR light

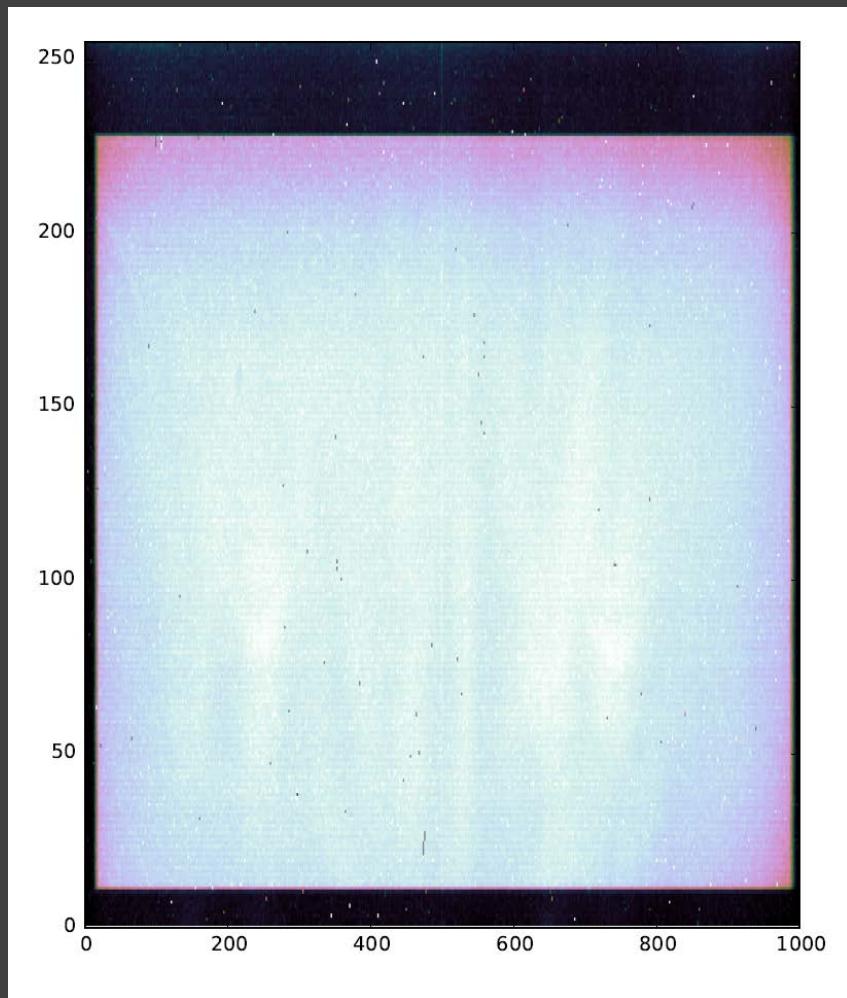


First SWIR light as seen in CSL

# New OGSE - Black body (solidifying silver cell)

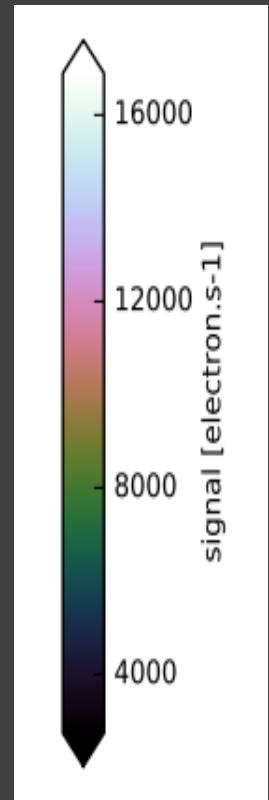


# Silver Cell Black Body – Absolute (Ir)Radiance



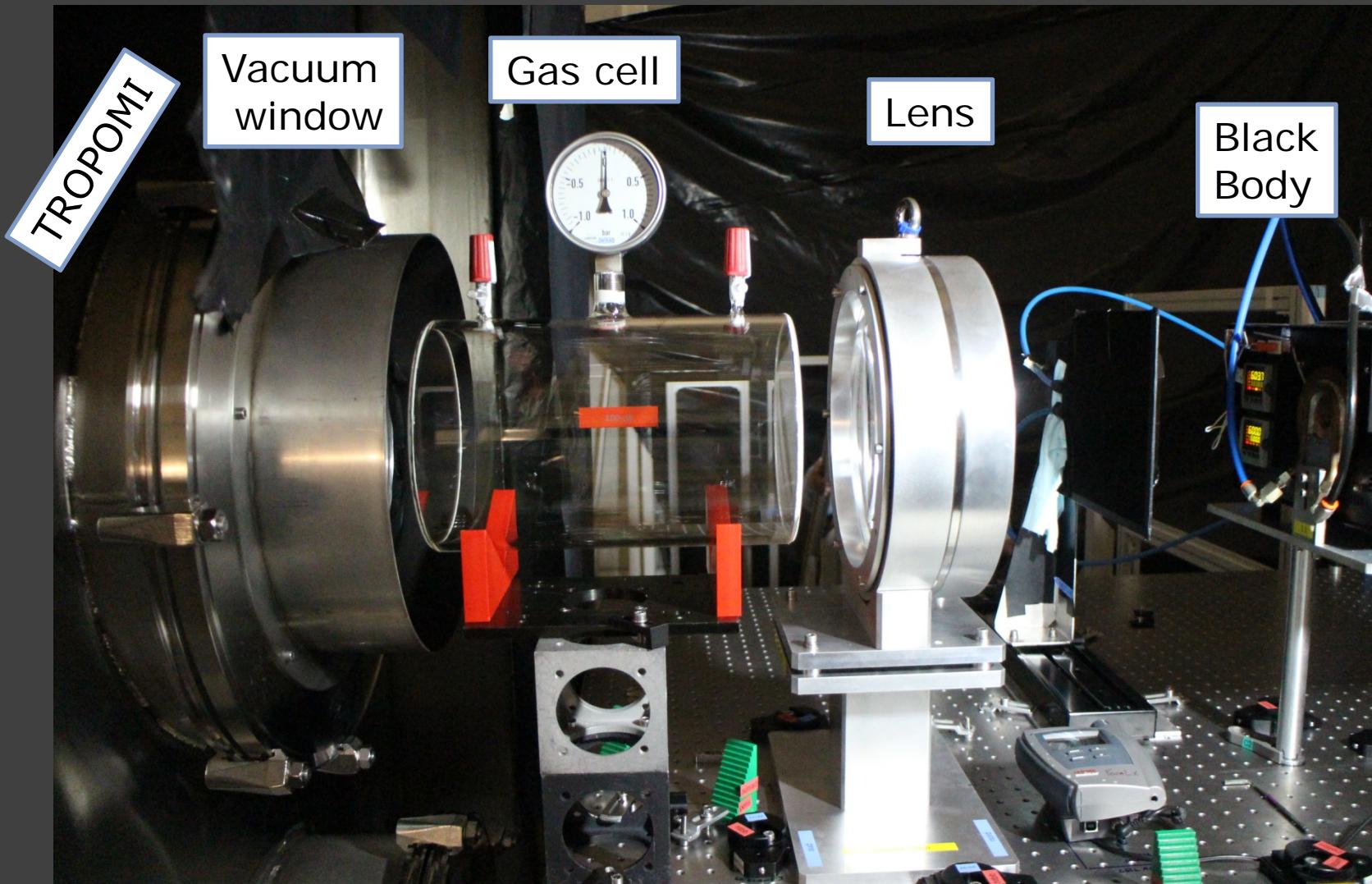
Wavelength (2300-2390nm)

109 degrees viewing angles



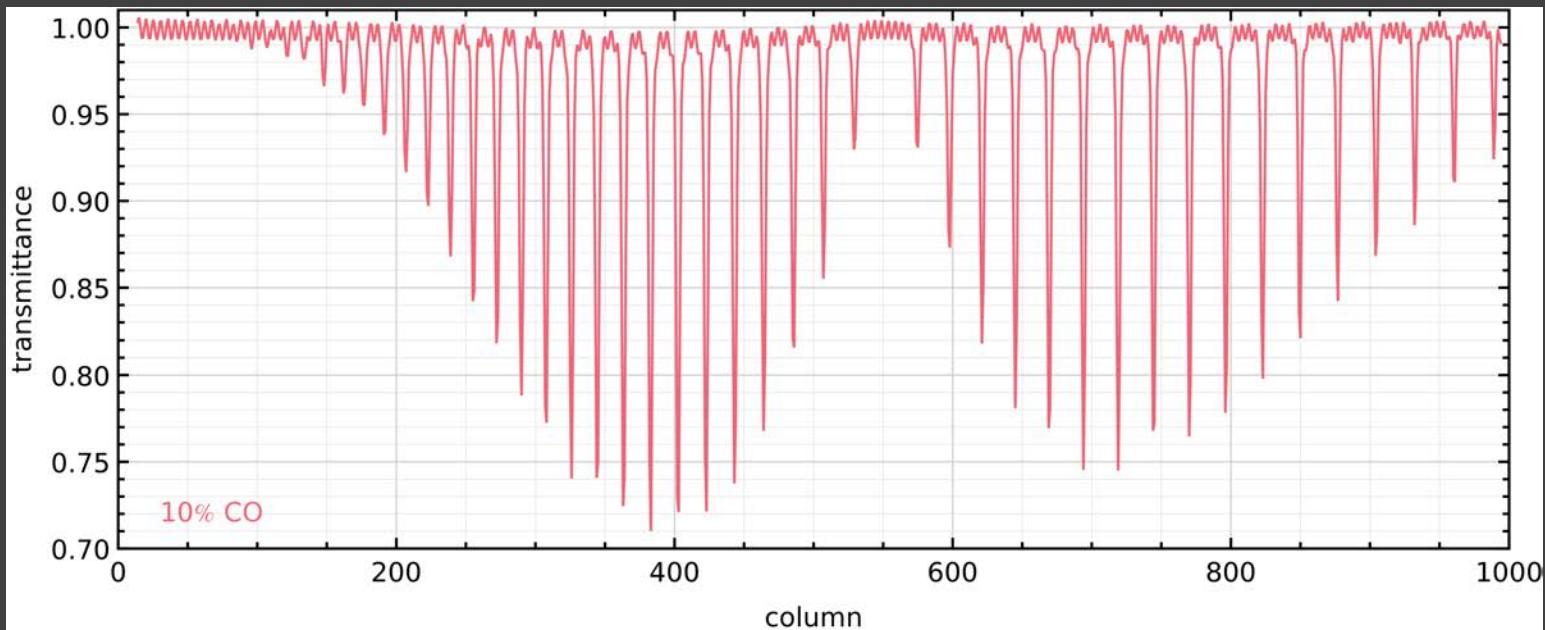
Preliminary analysis:  
0.1% stability

# Wavelength calibration



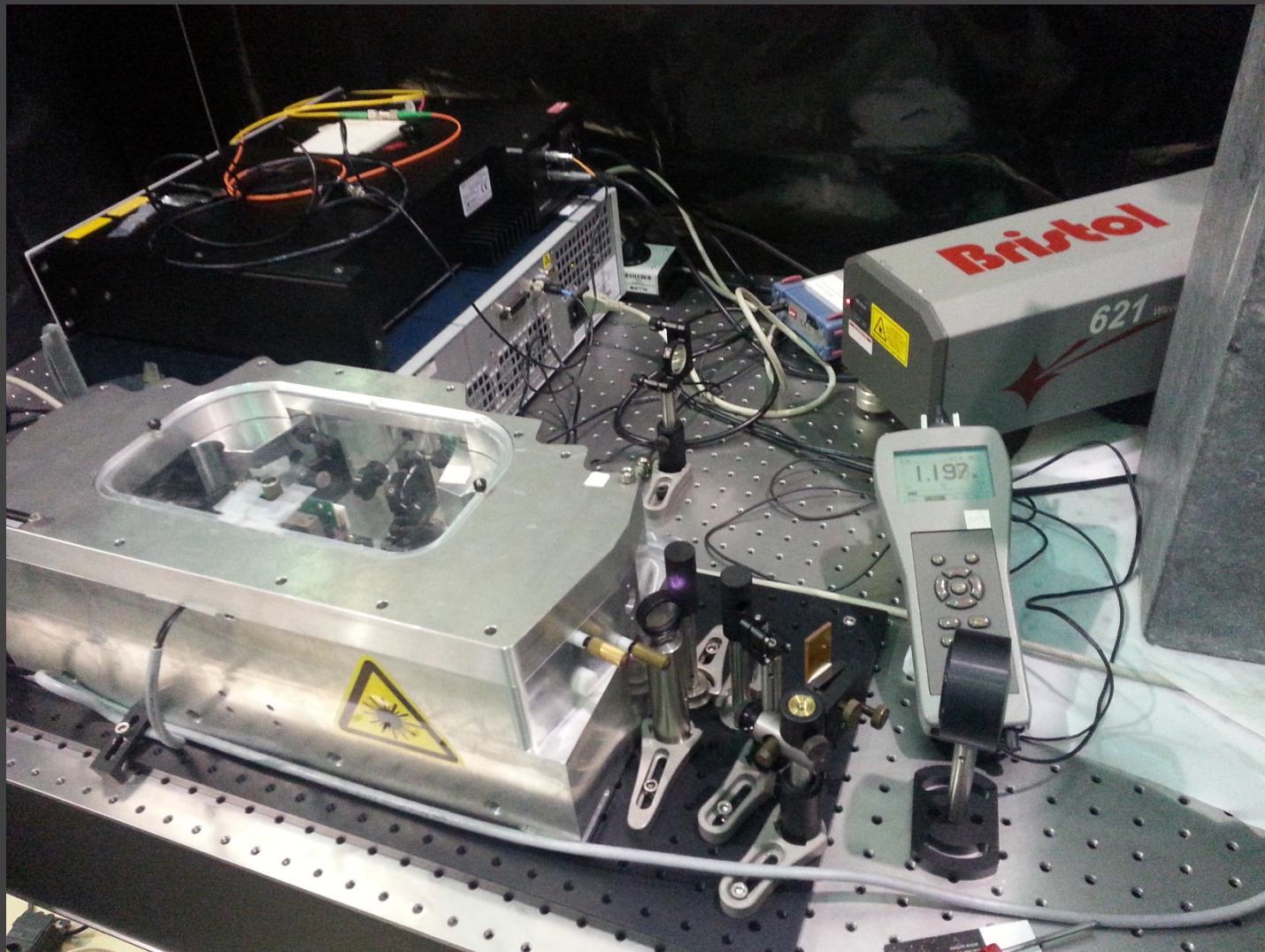
# Wavelength calibration

CO 10% Gas cell



Uncertainty after validation 0.003 nm (0.03px)

# SWIR custom-made tunable laser (VSL)

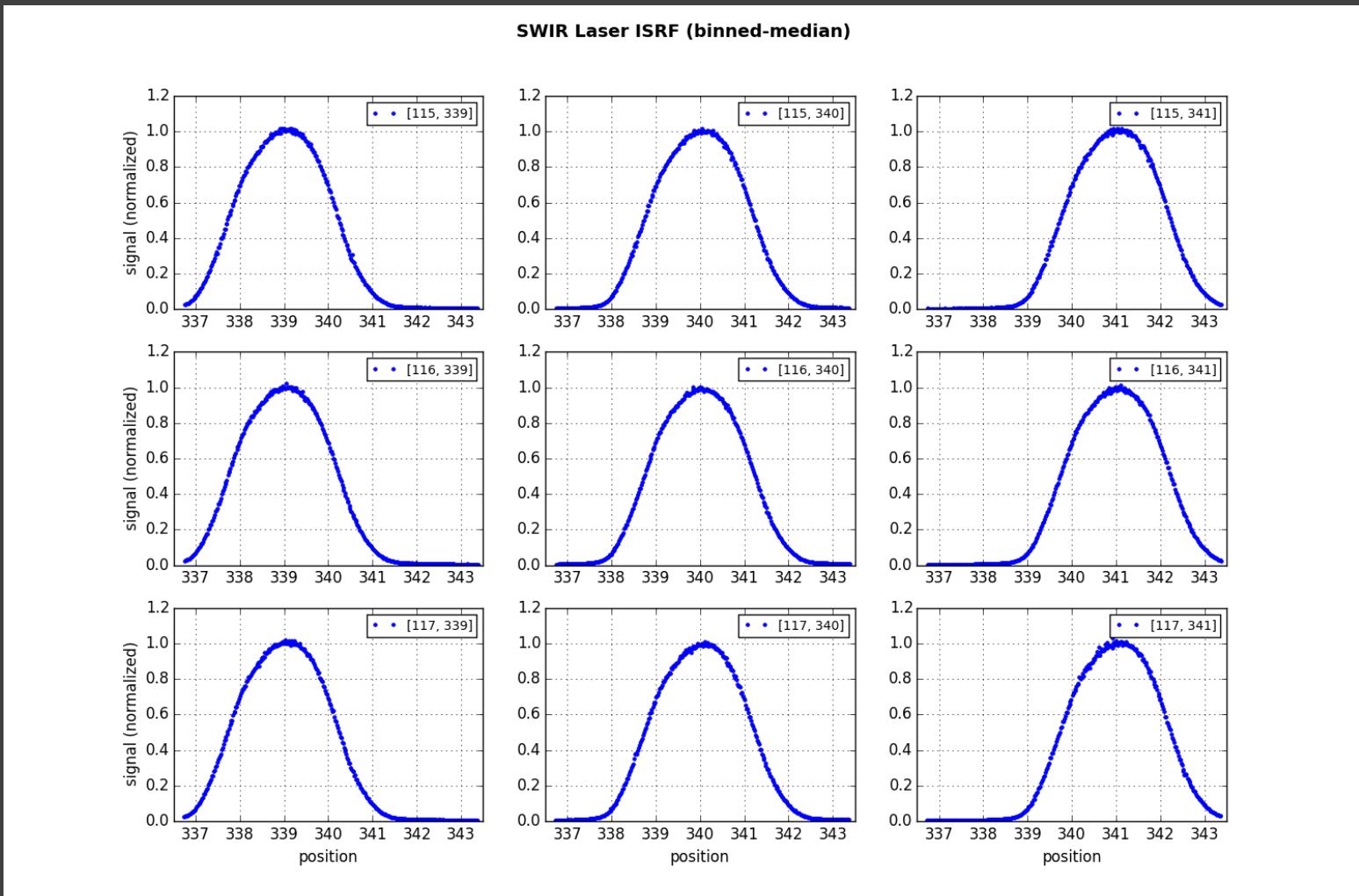


High  
Power  
Needed:  
2Watt

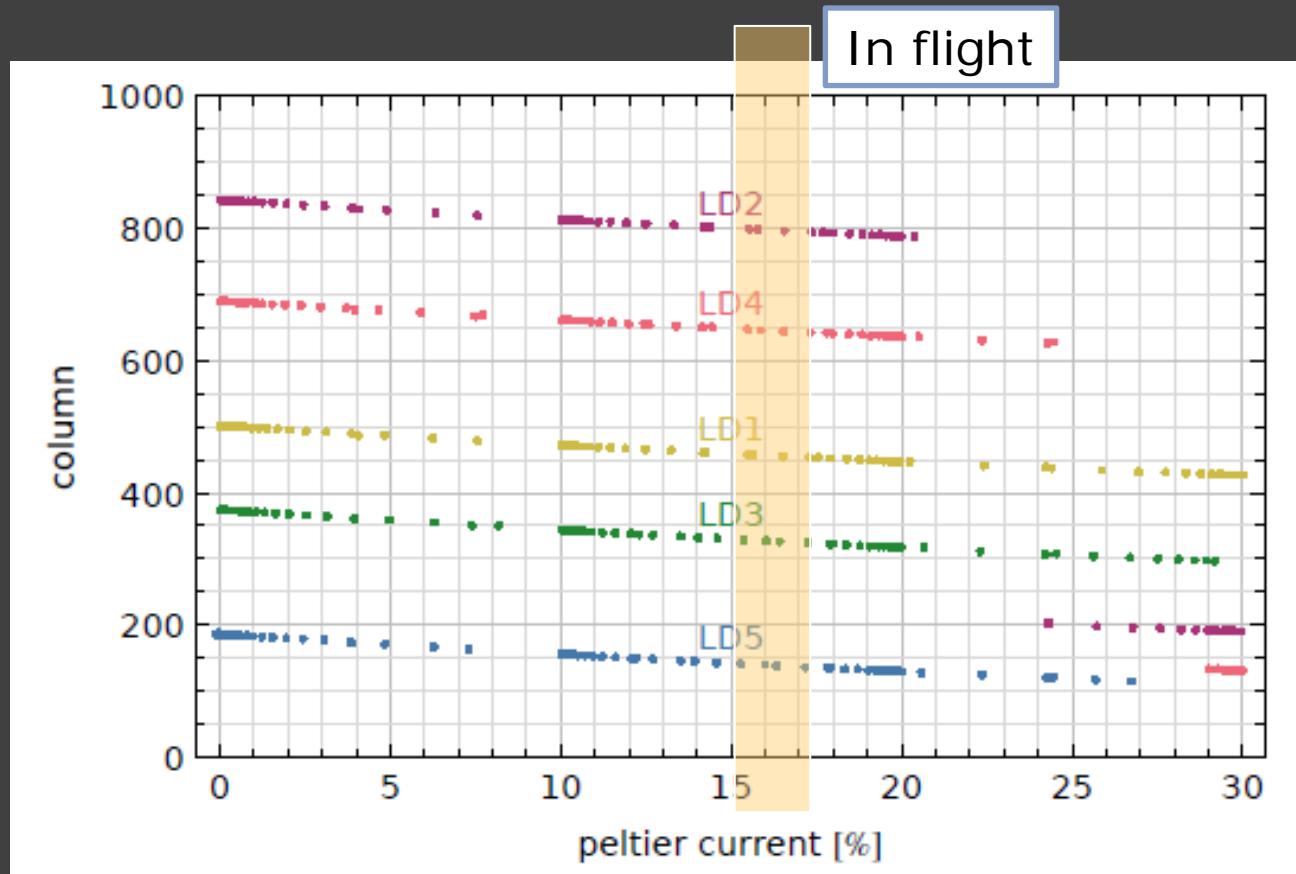
# Instrument Spectral Response Function

Scan laser over both wavelength and angle

0.002nm steps

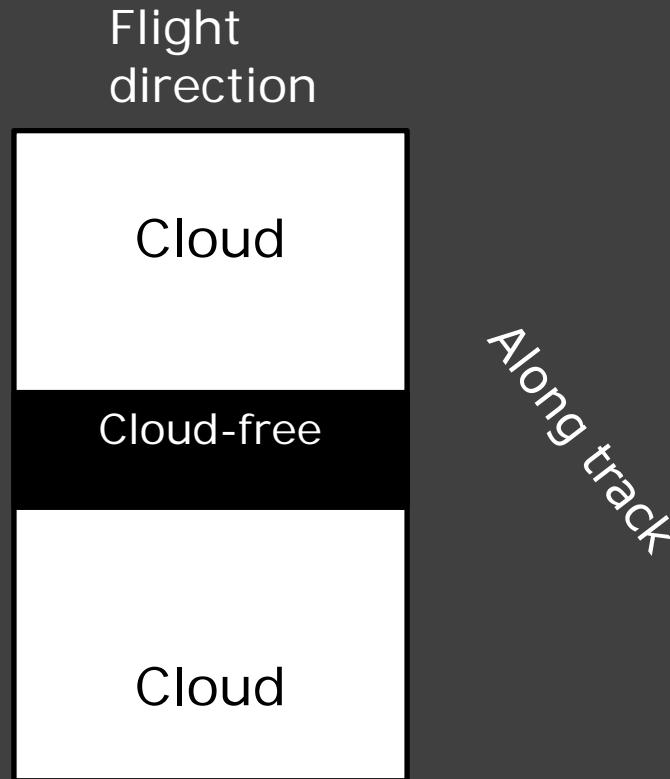


# ISRF: On-board laser diodes characterization



5 Internal Laser Diodes with variable Peltier-cooler power,  
and hence variable wavelength  
This will allow in-flight ISRF monitoring at 0.002nm resolution

# Straylight: Why should you care

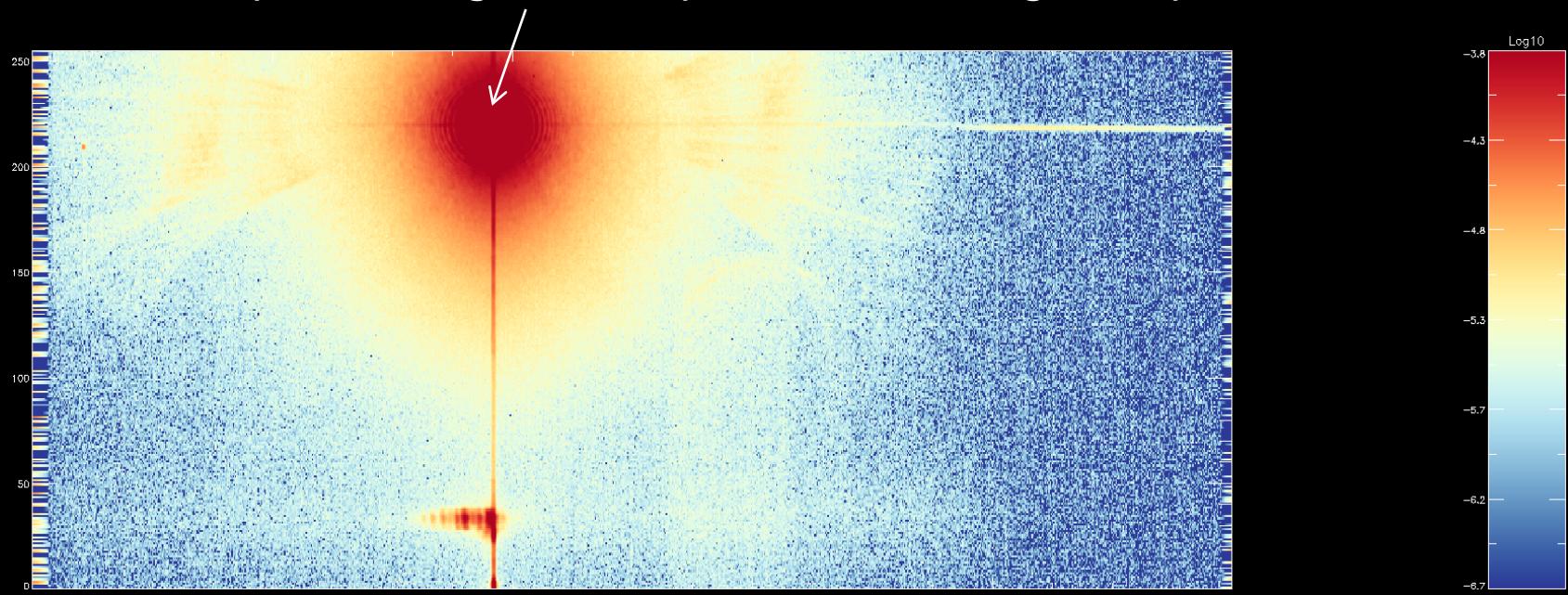


# Straylight

Cloud

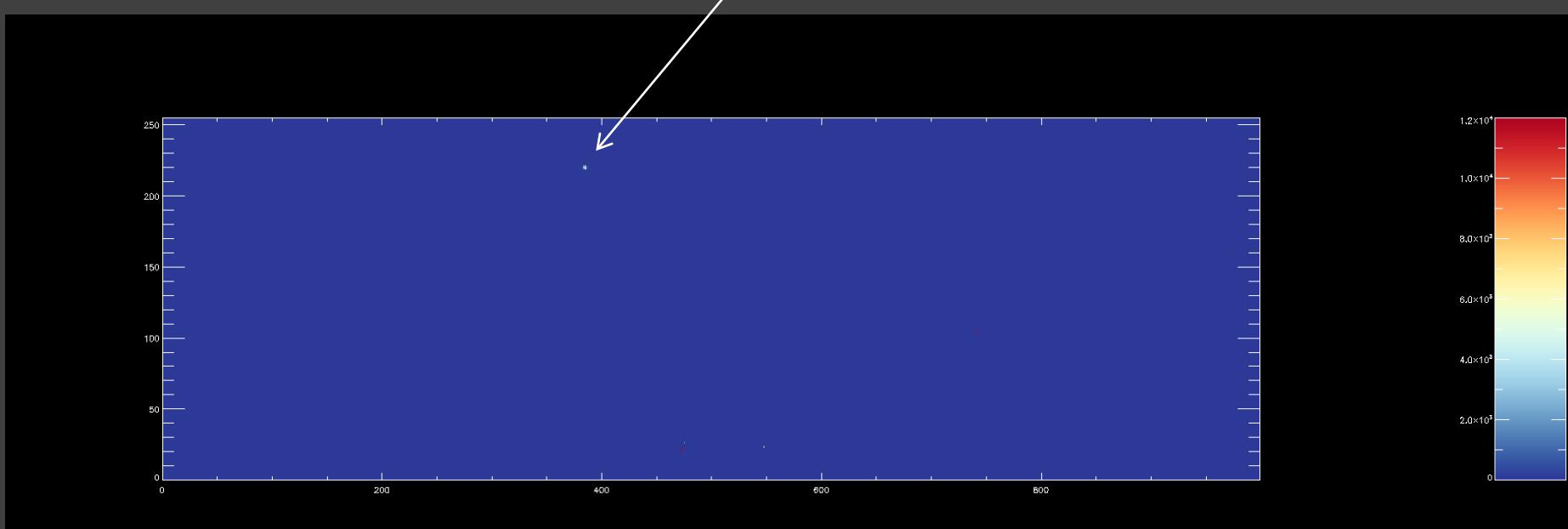
Area of interest

Laser under specific angle with specific wavelength, repeat for all

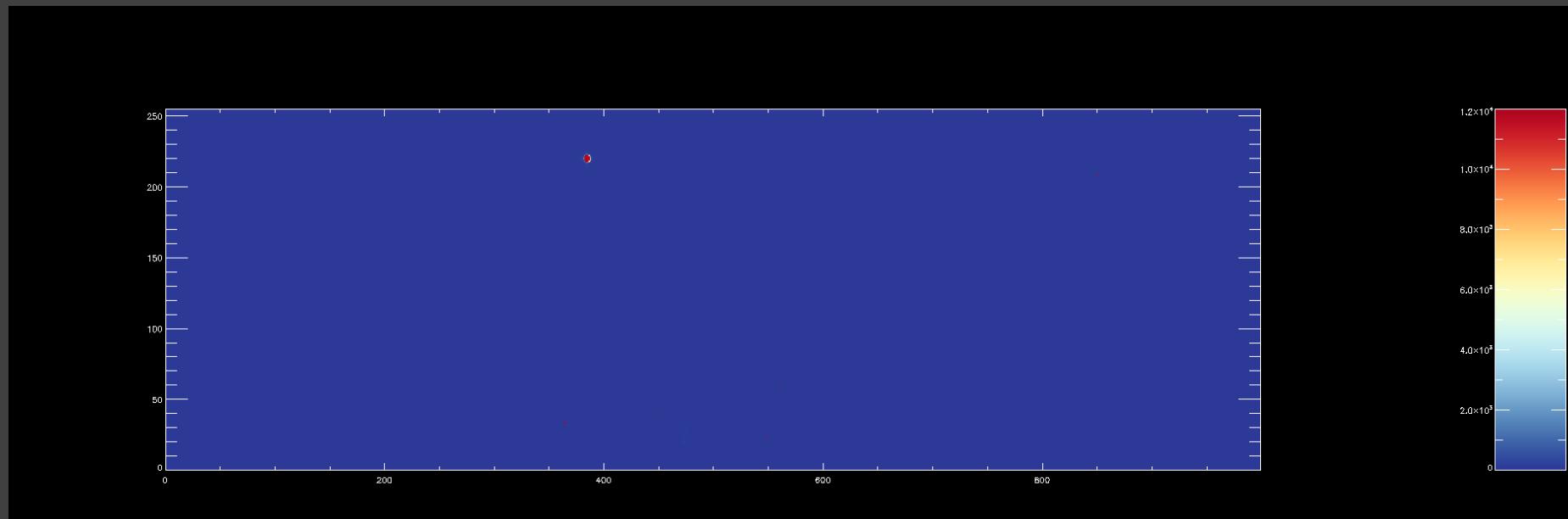


New method: varying integration times

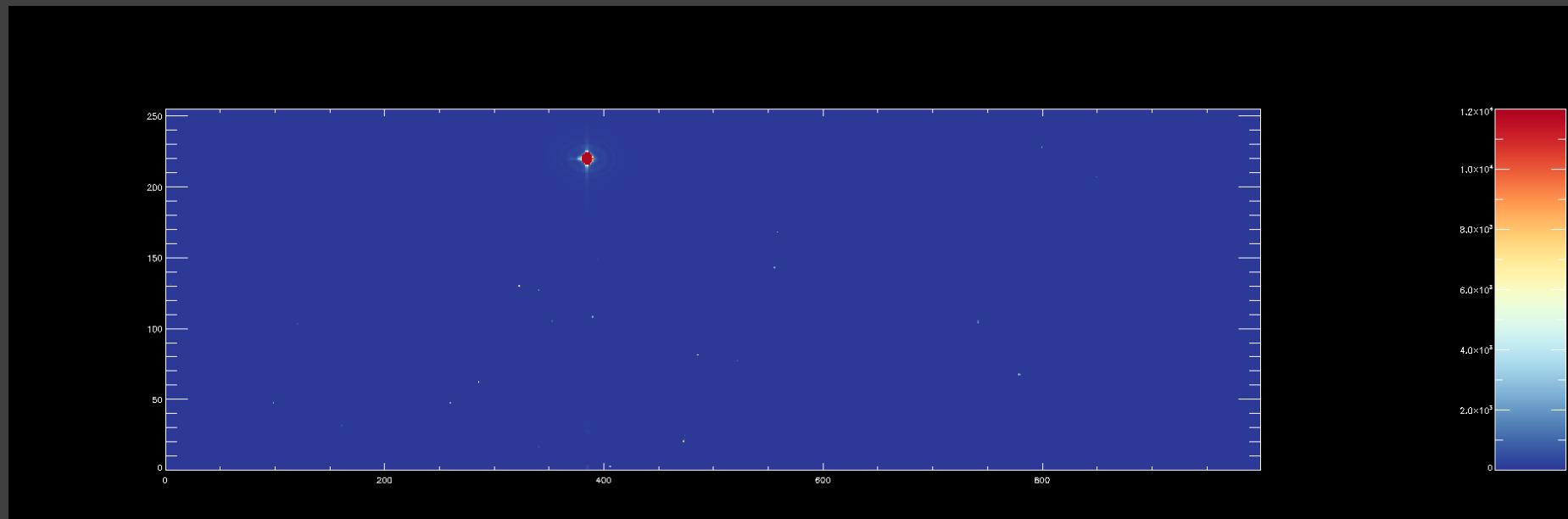
# Laser 0.2ms (remember that high power)



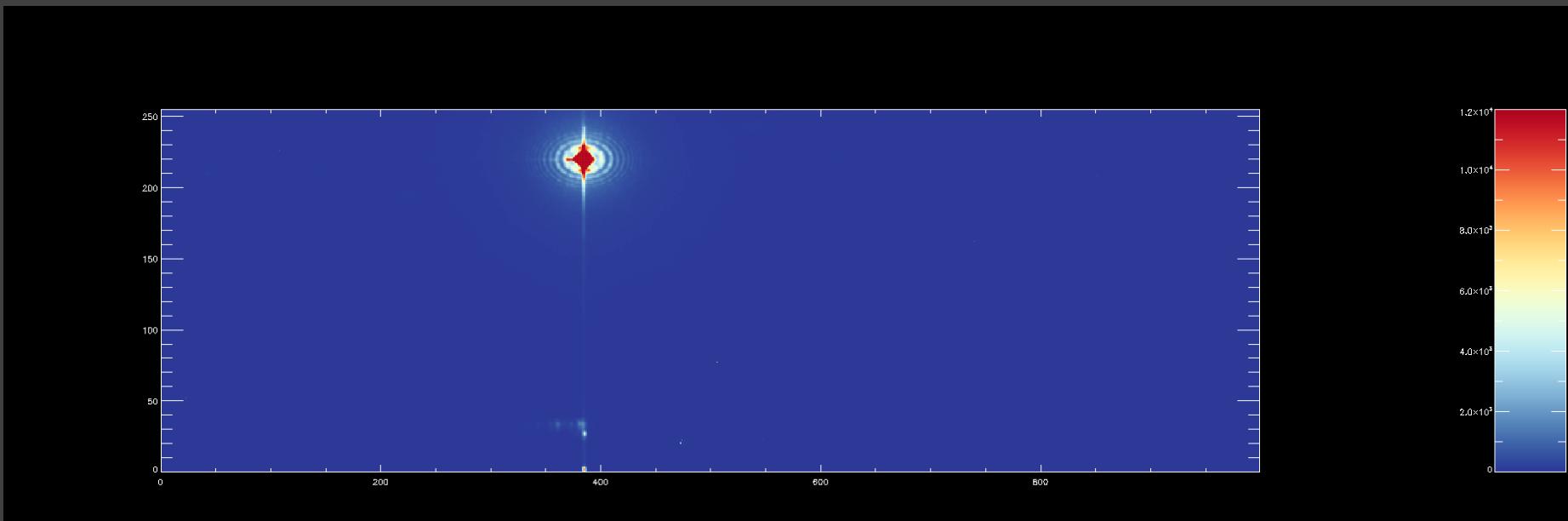
# Laser 4ms



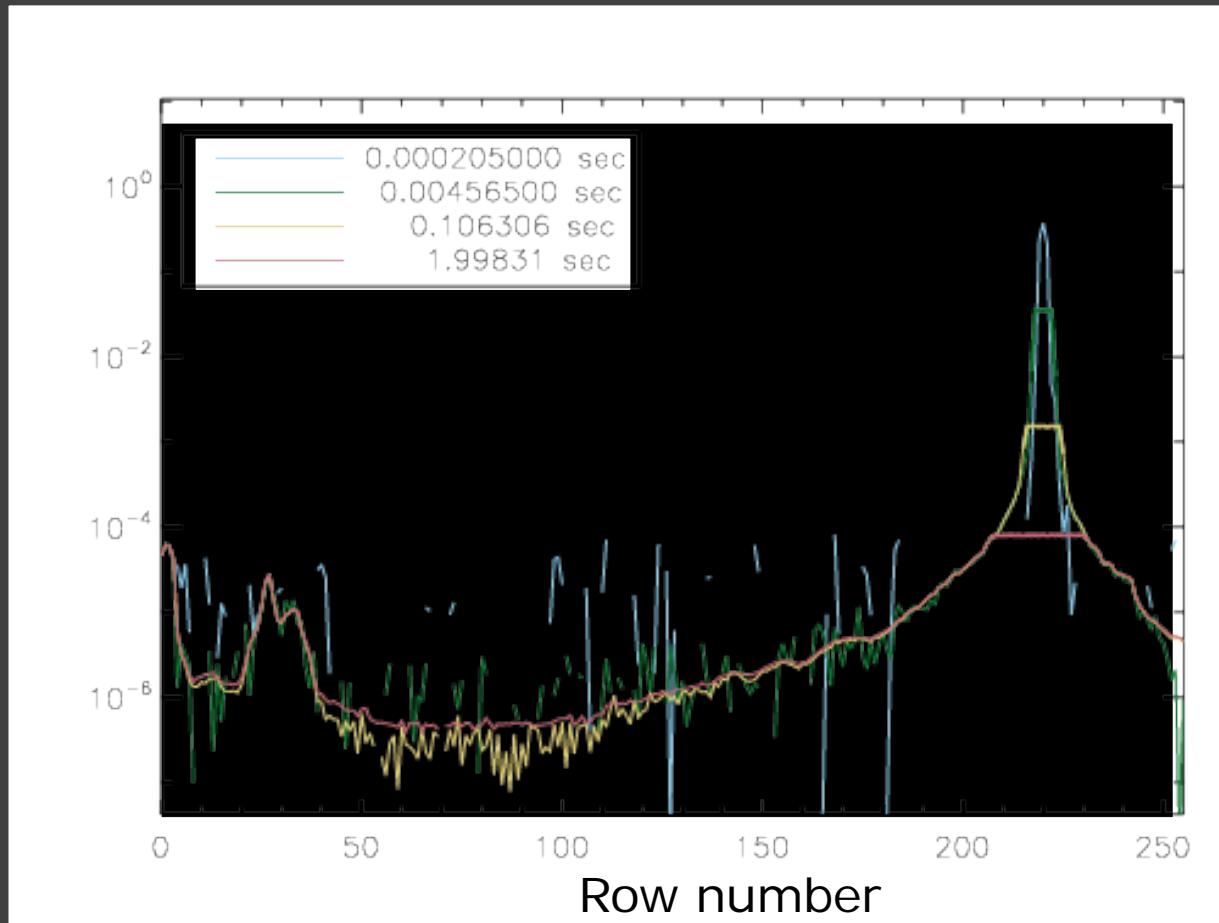
# Laser 100ms



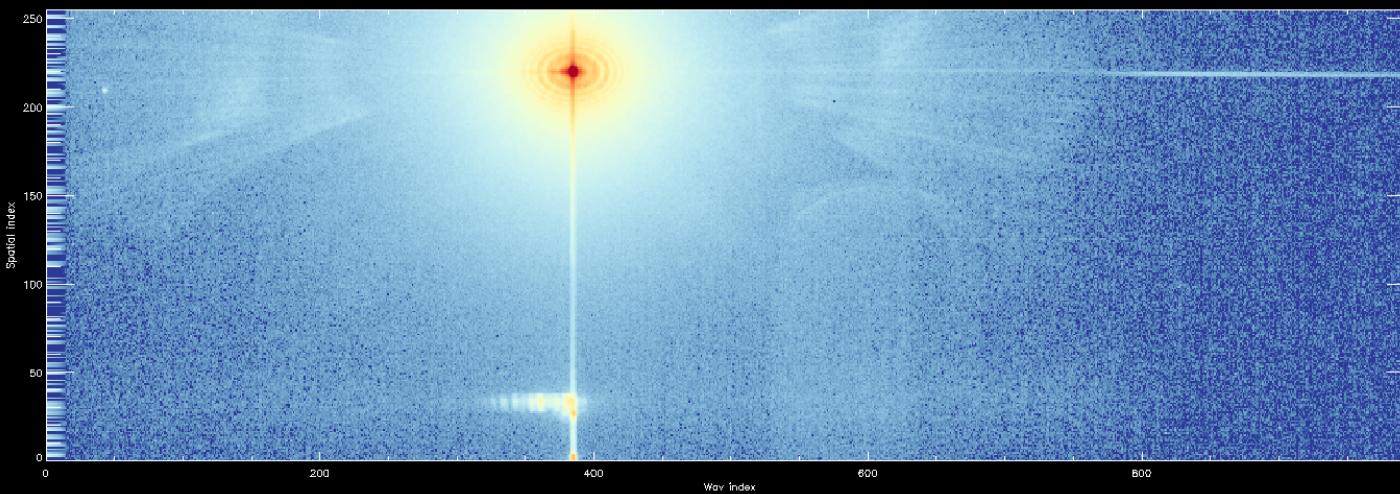
# Laser 2sec



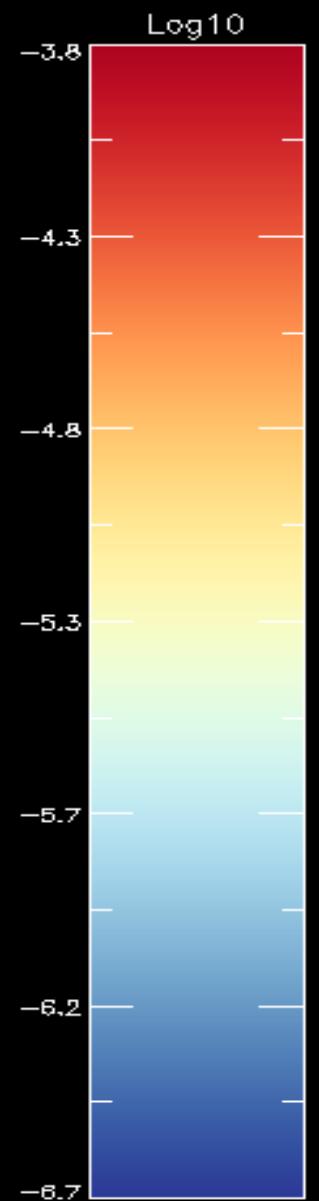
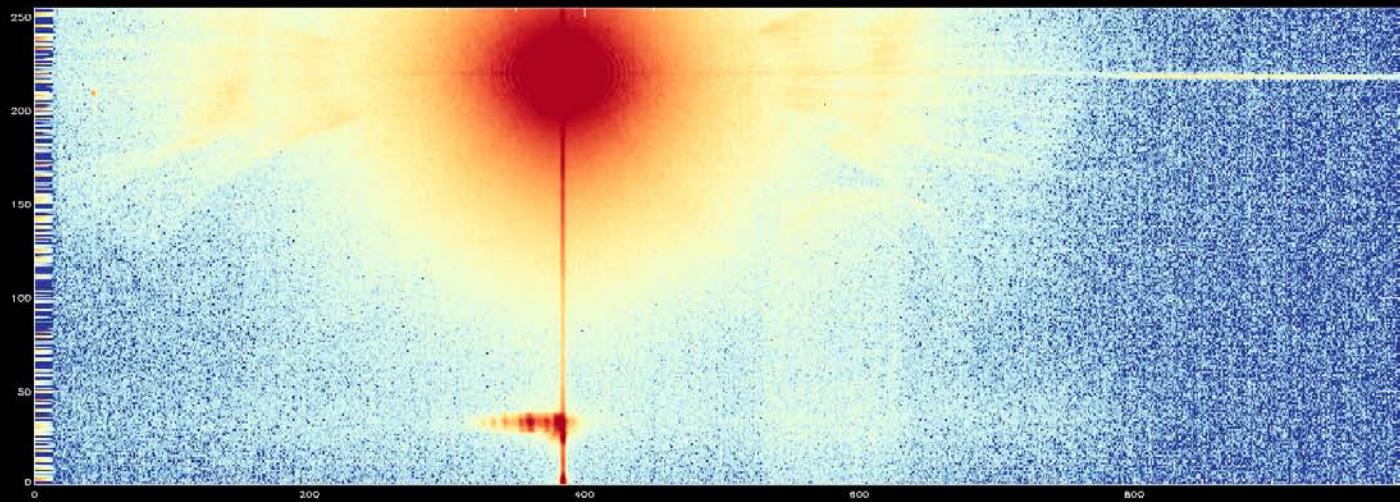
# Add it all together and stir



# Stray Light Laser combined (logarithmic)



Straylight measured to  $10e-6 \sim 10e-7$  or .001%  
Now to implement correction algorithm...



# Summary

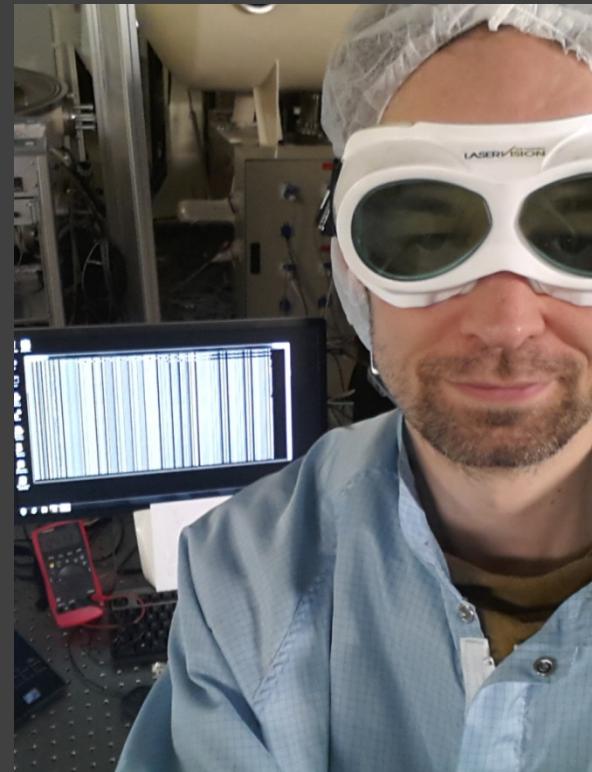
- Straylight
  - at 0.001% measured  
(allows 2 orders of magnitude reduction)
- Wavelength
  - 0.003 nm uncertainty
- Instrument Spectral Response function
  - for every pixel
  - at 0.002nm
- Laser Diodes in-flight monitoring
  - at 0.002nm
  - (for select few pixels)

Measurements meet requirements!

**Analyses on-going**

**Thank you**

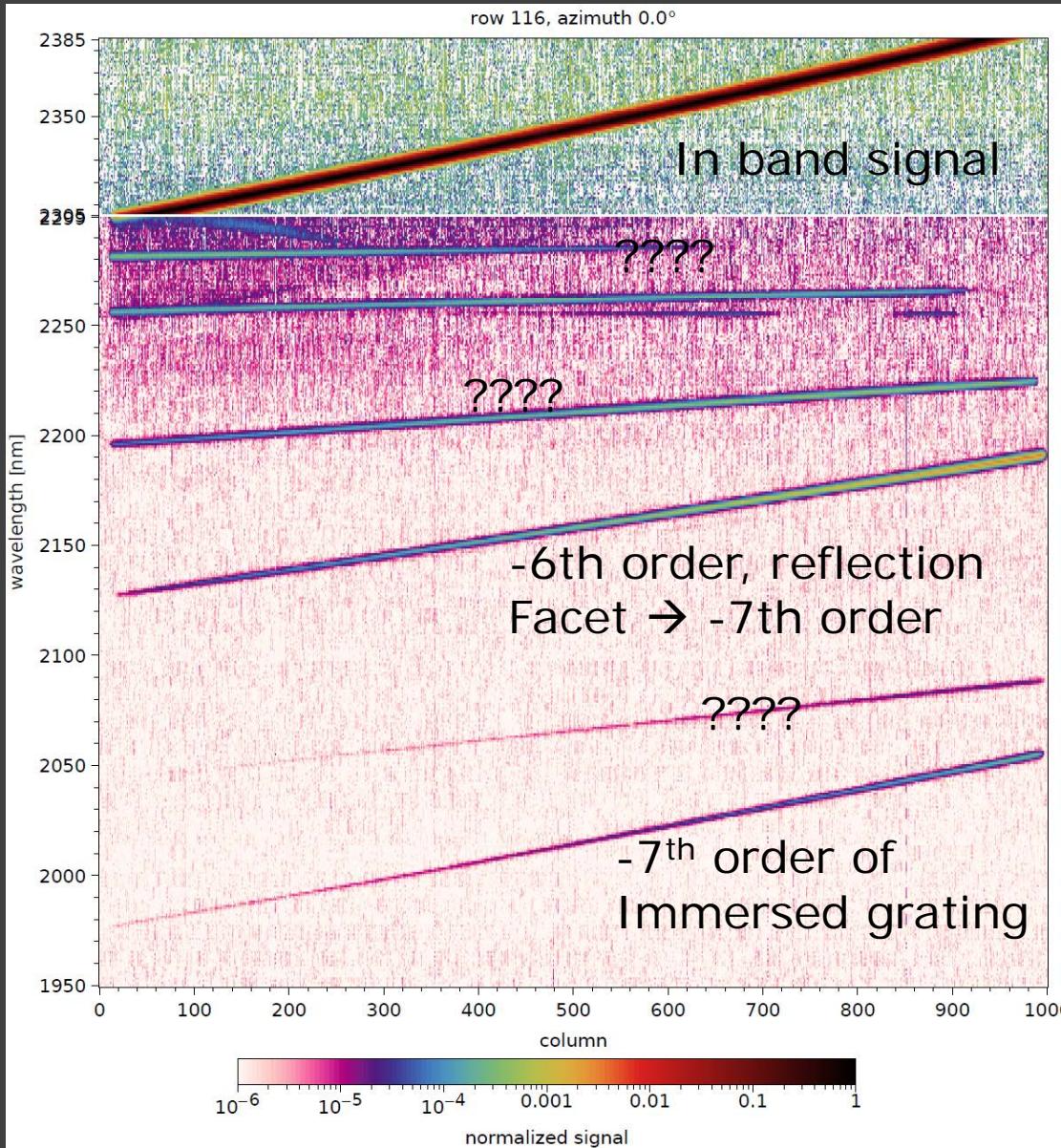
Your author Matthijs Krijger in action





# Backup Slides

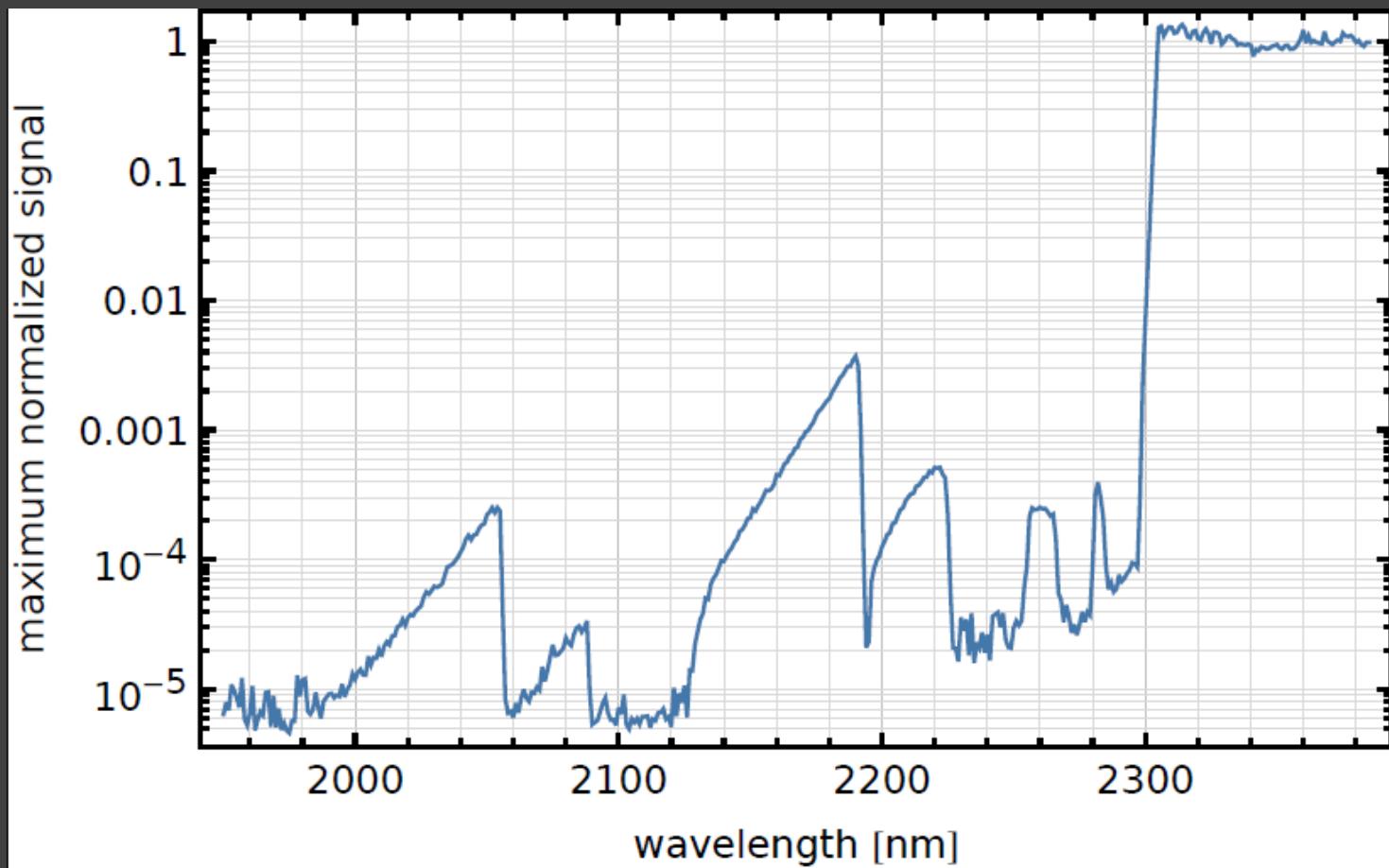
# Straylight Out of Band



Commercial  
Expla laser

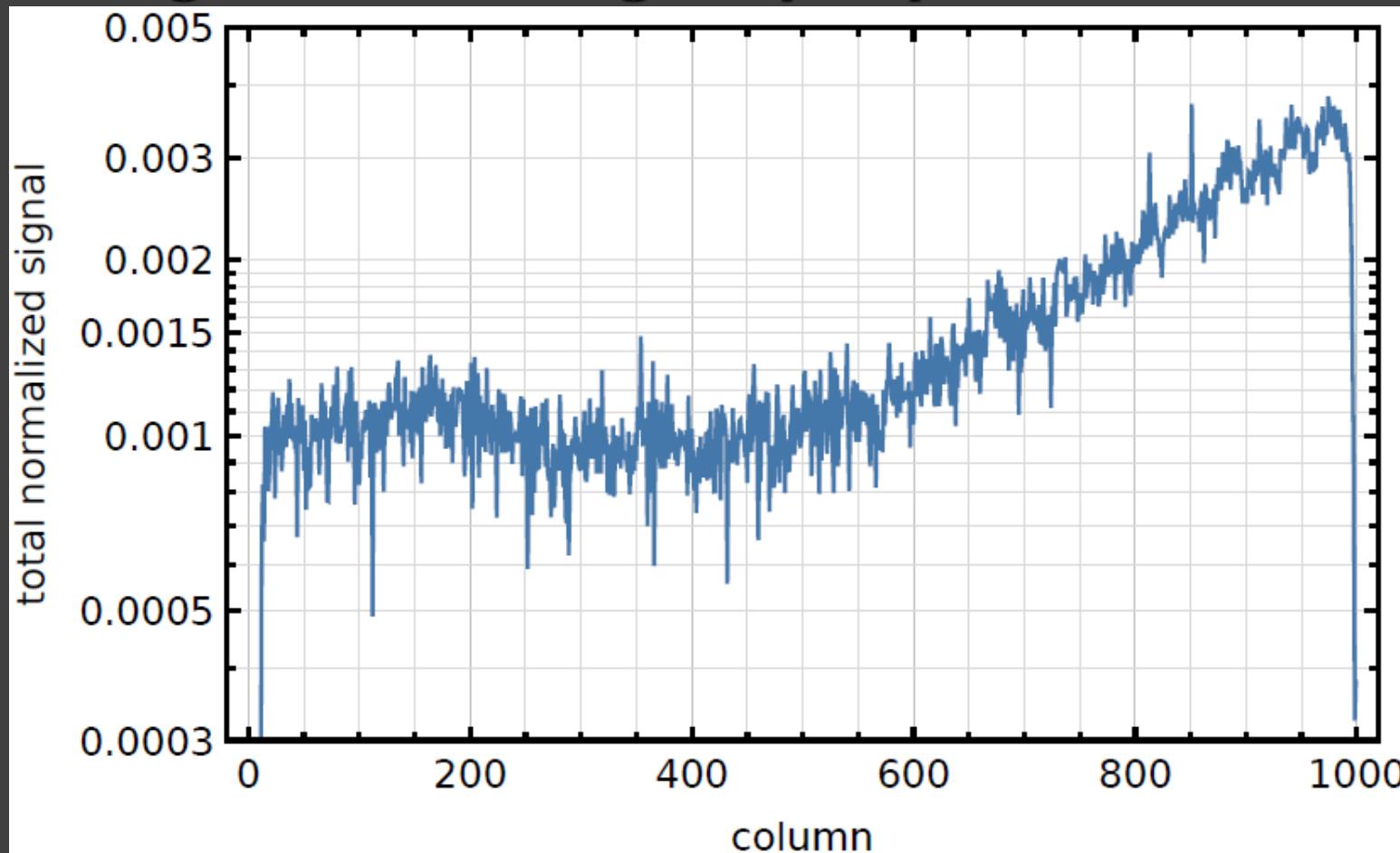
Several  
unexplained  
ghosts

# Maximum signal per wavelength



Maximum OOB straylight at 2190 nm: 0.35%.

# Integrated OOB signal per pixel

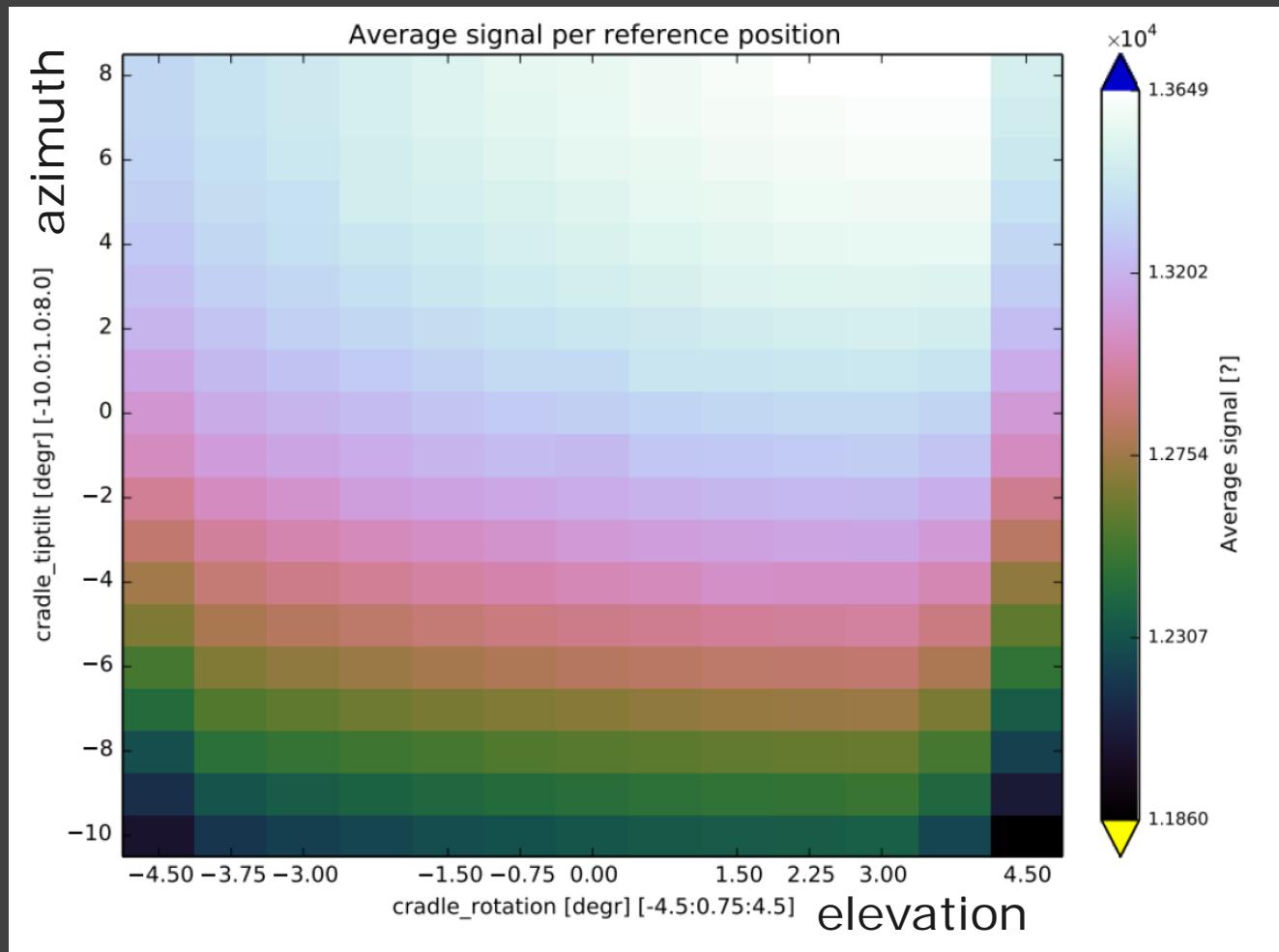


OOB straylight around 0.1%, up to 0.35% at the end of the band  
Measured up to 10e-3 or .001

# On ground calibration plan

category	session type	source	port	cradle	pattern
SWIR only	SWIRLS_EARTH_STRAY	N 		↔	
	SWIRLS_EARTH_ISRF	N 		↔	
	SWIRLS_EARTH_ISRF_INHOMO	N 		↔	
	FPBB_EARTH			↔	
	CELL_EARTH	 		↔	
	CELL_EARTH_INHOMO	 		↔	
	SWIRLS_SUN_STRAY	N 		○	
	SWIRLS_SUN_ISRF	N 		○	
	FPBB_SUN			○	
	FPBB_SUN_BSDF			○	
	CELL_SUN	 		○	
	LASER_DIODES	[  ]		○	
	LD_STRAY	[  ]		○	
	DLEDLIN_SW	[  ]		○	
	WLSLIN_GO	[  ]		○	

# Relative Irradiation sensitivity



Rotating TROPOMI

Seasonal angle

Sunrise angle