

SEOM – INSARAP: Sentinel-1 InSAR Performance Study with TOPS Data

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Special Considerations in the TOPS Case

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Knowledge for Tomorrow

Coregistration and Azimuth Shifts*

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Interferometry with TOPS: coregistration and azimuth shifts

EUSAR 2014, Berlin, Germany.



Impulse response and interferometric phase

- Impulse response (zero-Doppler focusing geometry)

$$s(\tau, t; r_0) = s_r(\tau - \tau_0) \cdot s_a(t - t_0 - k \cdot (\tau - \tau_0)) \cdot \exp\left[-j \frac{4\pi}{\lambda} r_0\right] \\ \cdot \exp[j2\pi f_0 \cdot (\cos \beta - 1) \cdot (\tau - \tau_0)] \cdot \exp[j2\pi f_{dc} \cdot (t - t_0)]$$

True range position of scatterer within resolution cell

- Interferometric phase for azimuth shifts

- $\Delta\phi = 2\pi \cdot f_{DC}(t) \cdot \Delta t$

- Variable Doppler centroid (TOPS)

- Constant shift => azimuth phase ramp in each burst
- Local shift => local phase contribution (bug or feature?)



One could estimate the azimuth shifts...

- Phase error due to azimuth mis-registration (in presence of Doppler centroid)

$$\sigma_{\varphi_{AZ}}^2 \approx 6 \left(\frac{f_{dc}}{B_d} \right)^2 \frac{1 - \gamma^2}{N \gamma^2}$$

- Azimuth coregistration performance depends on coherence, samples and Doppler bandwidth
 - The phase error depends on Doppler centroid (and coregistration error)
- Performance of (range) phase measurement, the well known CRB

$$\sigma_{\varphi_{IF}}^2 \approx \frac{1}{2} \frac{1 - \gamma^2}{N \gamma^2}$$

- On the same resolution (same N) the coregistration-related term will always dominate in the TOPS case! (as soon as the Doppler centroid is comparable to the bandwidth)



Two cases: systematic shifts & geophysical shifts

- Systematic shifts:
 - Orbital accuracy, timing errors, (Earth tides)*
 - Can be corrected to sub-centimetric accuracy
- Geophysical shifts:
 - The effect on the phase should only be correctly interpreted
 - Users (and their models) should be aware

*M. Eineder *et. al.*, "Imaging Geodesy – Toward Centimeter-Level Ranging Accuracy with TerraSAR-X," *IEEE Trans. on Geoscience and Remote Sensing*, vol. 49, no. 2, Feb. 2011.



Systematic azimuth shifts

- Systematic shifts:
 - Orbital accuracy, timing uncertainty
 - One single value for a long scene (tens of kilometers)
 - Lots of averaging opportunities
 - Many independent samples
 - Burst edges and overlaps (high sensitivity)
 - Elimination of outliers might be necessary
- Sub-centimetric accuracy (for Sentinel-1, 20m azimuth resolution)
 - Coregistration accuracy $\sigma_x = \sqrt{\frac{3}{2N} \frac{1-\gamma^2}{(\pi\gamma)^2}} \rho_{az} = 1\text{cm @ } 1000\text{ km}^2 (\gamma = 0.25)$
 - Resulting phase accuracy $\sigma_\varphi = 360 \frac{f_{dc}}{v} \sigma_x = 3\text{ deg } (f_{dc} = 6\text{kHz}, v=7000\text{ m/s})$
- Long wavelength signals in azimuth direction, e.g. Earth tides, will be absorbed in this correction (they end up in the shift information)



Geophysical shifts

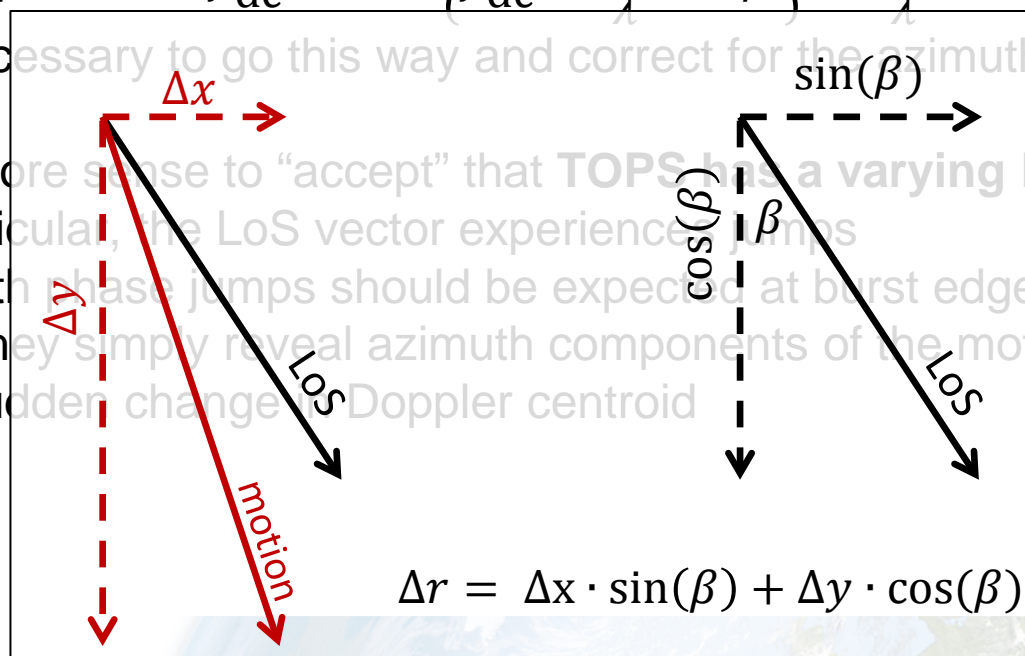
- It is not possible to compensate satisfactorily local shifts (at the scale of the interferometric product)
- The correction is really an attempt to separate azimuth and zero-Doppler component for a slightly squinted geometry

$$\Delta \phi = 2 \pi \cdot f_{dc} \cdot \Delta t = \left\{ f_{dc} = \frac{2v}{\lambda} \sin \beta \right\} = \frac{4\pi}{\lambda} \cdot \Delta x \cdot \sin \beta$$

- But is it necessary to go this way and correct for the azimuth LoS component?

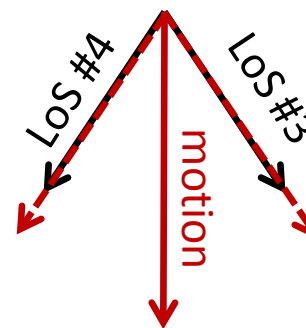
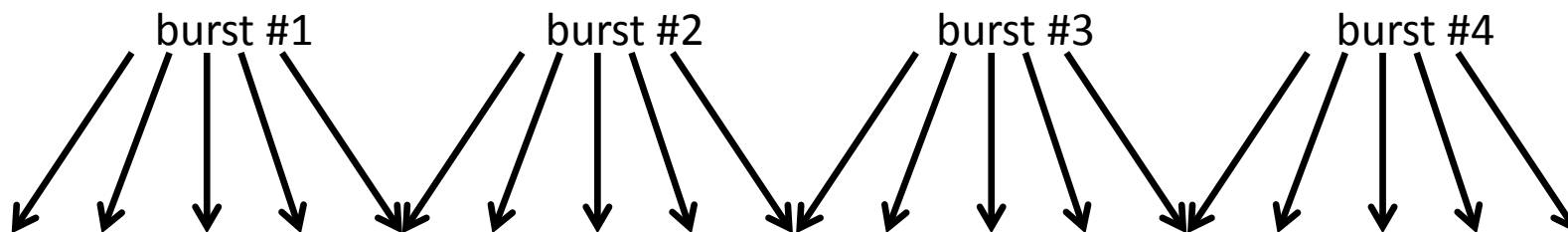
- It makes more sense to “accept” that TOPS has a varying line of sight!

- In particular, the LoS vector experiences jumps
- Azimuth phase jumps should be expected at burst edges
- They simply reveal azimuth components of the motion, sensed by a sudden change in Doppler centroid

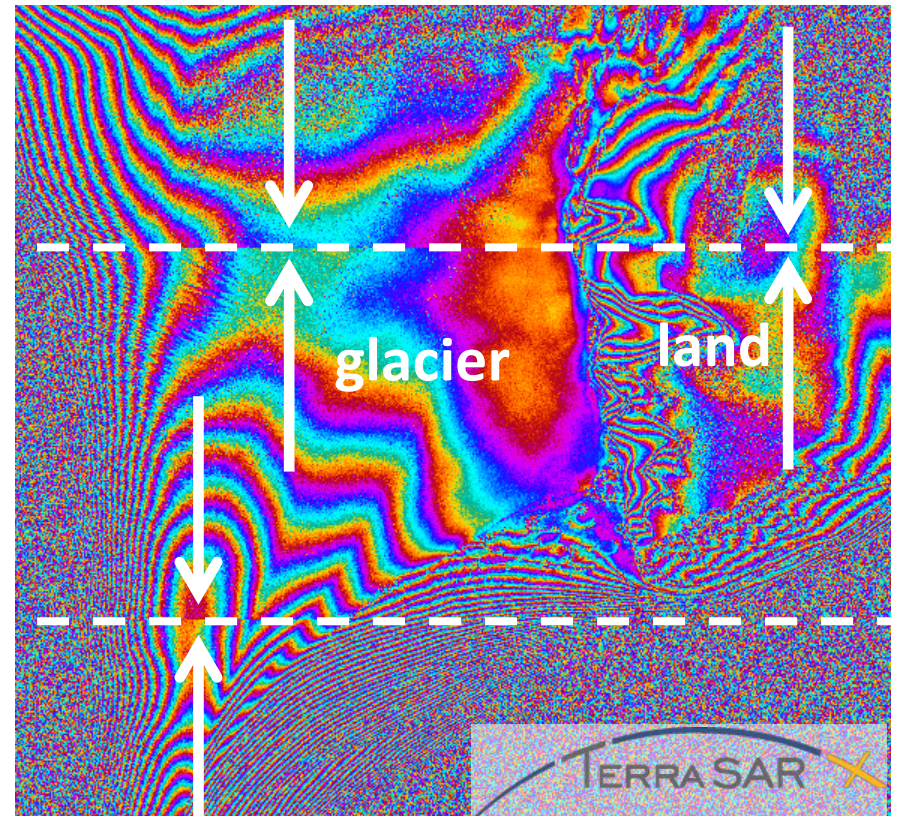
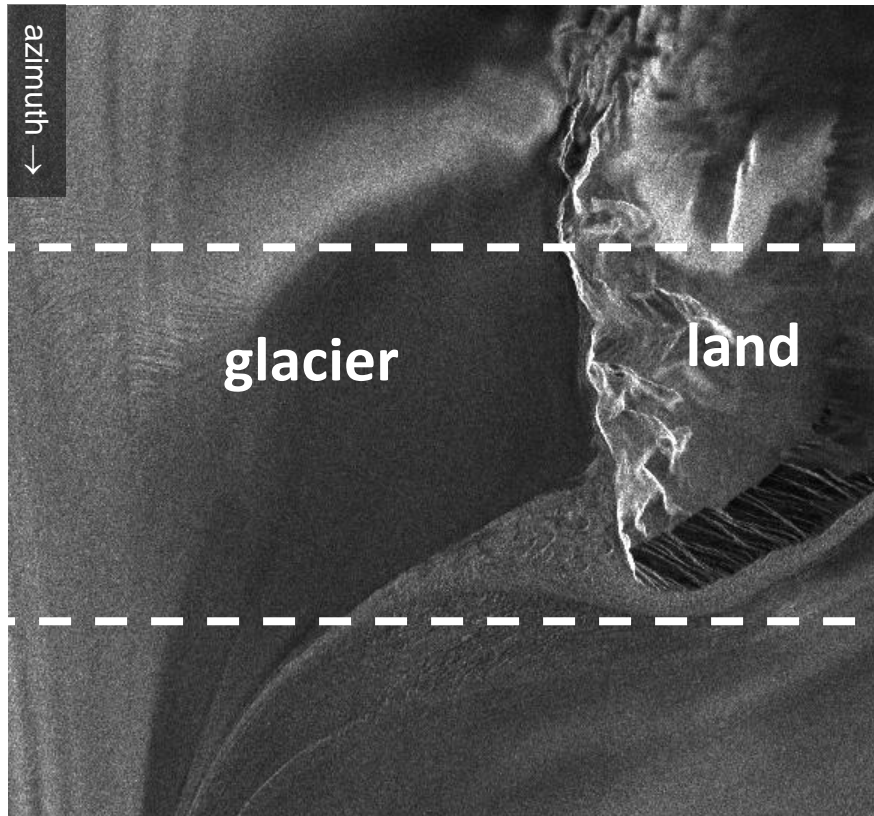


Variation of line of sight along azimuth

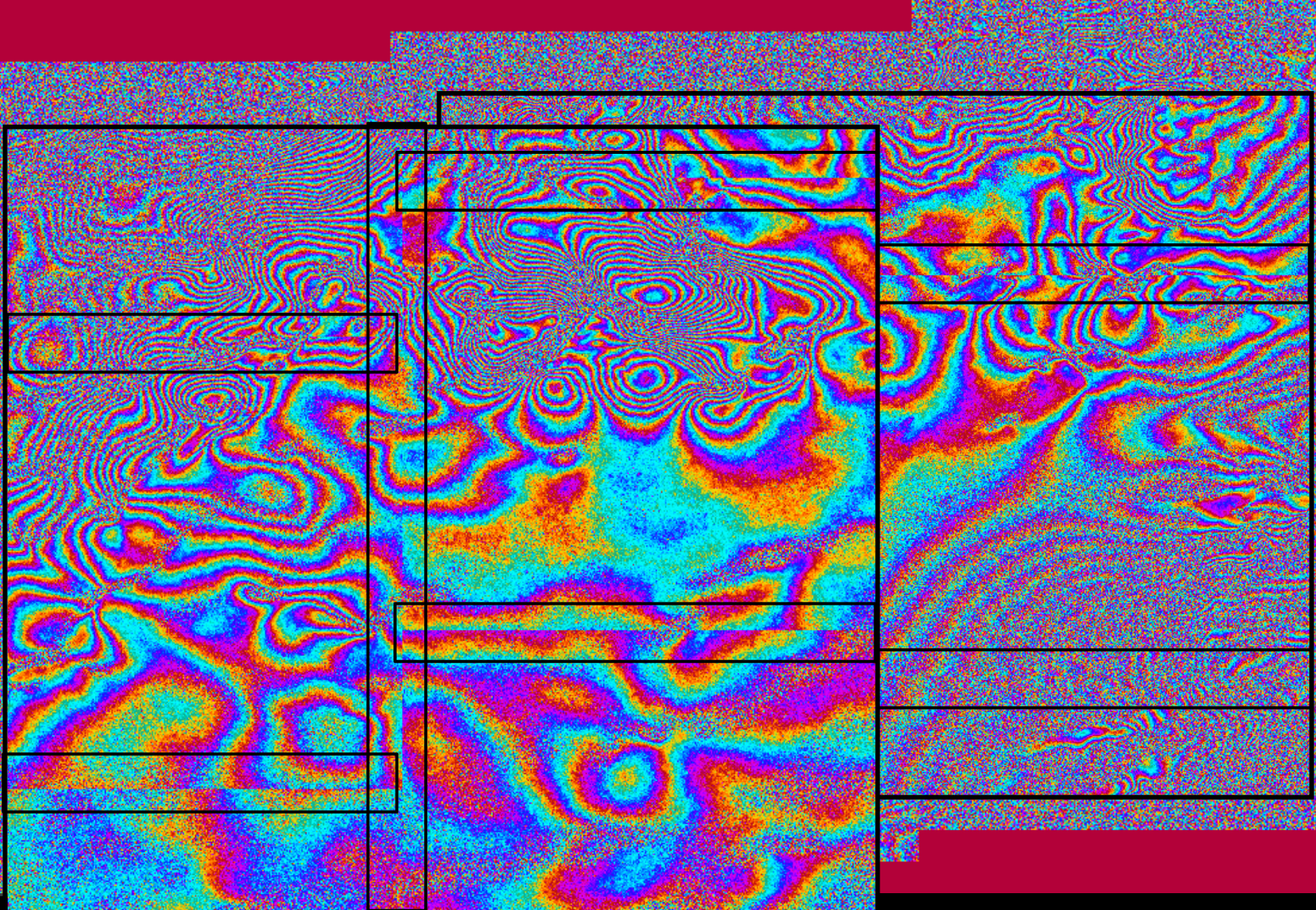
- SAR interferometry is sensitive to motion in line of sight
- Discontinuities in LoS are likely to appear in the interferogram
- Jumps in the interferograms will be normal (small or large)



Example With TerraSAR-X Data



Phase Discontinuities over Pine Island Glacier



azimuth →

Proposed methodology

- Correct with Enhanced Spectral Diversity for a constant shift or slow varying term (geolocation/timing error)
 - Burst-periodic phase ramps should disappear
 - Optionally, assume global offset is small enough to be ignored (with precise orbits)
- For phase variations or jumps due to local azimuth shifts
 - Models / users should be aware of the azimuth-varying line of sight
 - They should take care in phase unwrapping if the phase discontinuity is significant
 - Estimating the local azimuth shift, coregistering, unwrapping, reinserting the removed phase
 - Unwrapping each burst independently and solving for phase ambiguity, e.g. with the help of radargrammetry
 - Only large azimuth shifts will affect phase unwrapping: 75 cm (azimuth) → 180 deg phase jump. Earthquakes, glaciers...



Ionospheric Scintillations & TOPS



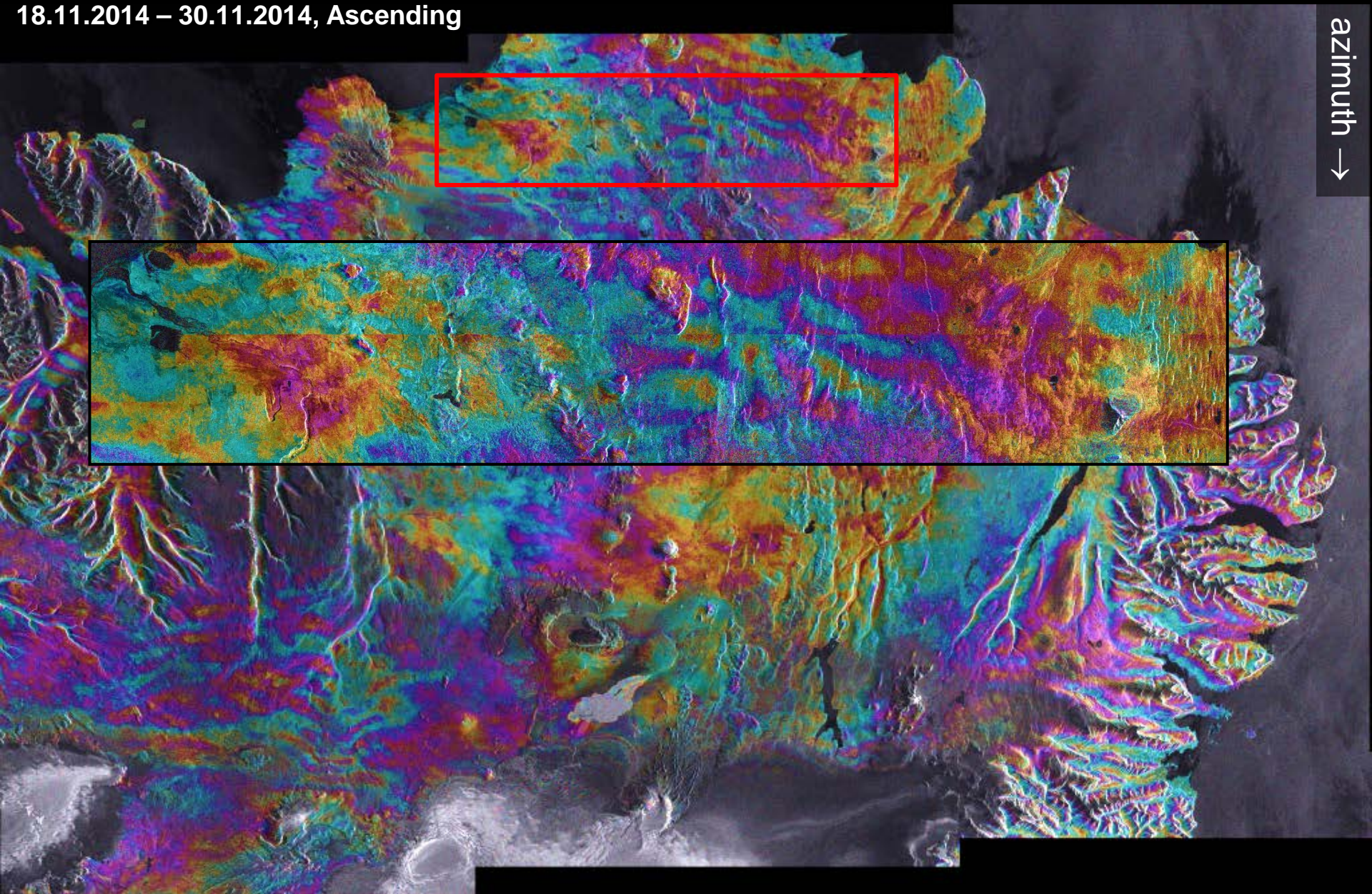
Ionospheric Scintillations and TOPS

- Ionospheric scintillations are specially intense and frequent close to the equator, but also at higher latitudes.
- Lower frequency bands are more affected.
- They can introduce azimuth shifts locally up to decimeters at C-band.
- For TOPS, the azimuth shifts will turn into phase jumps at burst edges.
- In the following an example with Sentinel-1 data over Iceland is shown.



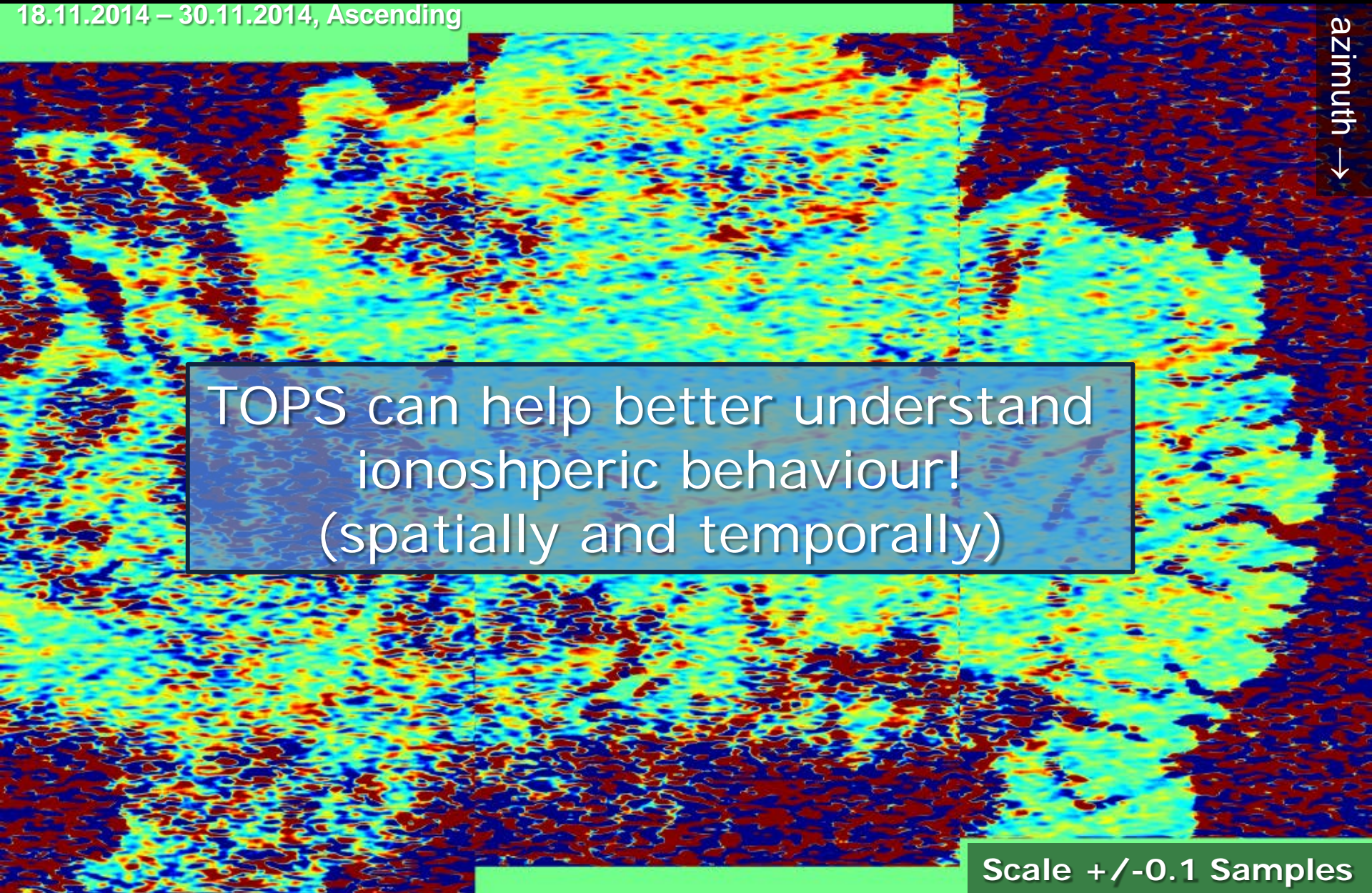
Iceland

18.11.2014 – 30.11.2014, Ascending



Iceland: Estimated azimuth offsets

18.11.2014 – 30.11.2014, Ascending



TOPS can help better understand ionospheric behaviour!
(spatially and temporally)

azimuth →

Scale +/-0.1 Samples

Speckle Tracking with TOPS Data*

*R. Scheiber, M. Jäger, P. Prats-Iraola, F. De Zan, D. Geudtner

Speckle Tracking and Interferometric Processing of TerraSAR-X TOPS Data for Mapping Nonstationary Scenarios

IEEE Journal of Selected Topics in Applied Earth Observations and Remote Sensing, online, early access.



Speckle Tracking with TOPS

- Large azimuth offsets cause spectral decorrelation

$$\Delta f_{shift} = \frac{K_a K_{rot}}{K_a - K_{rot}} \Delta_{az} / v_g$$

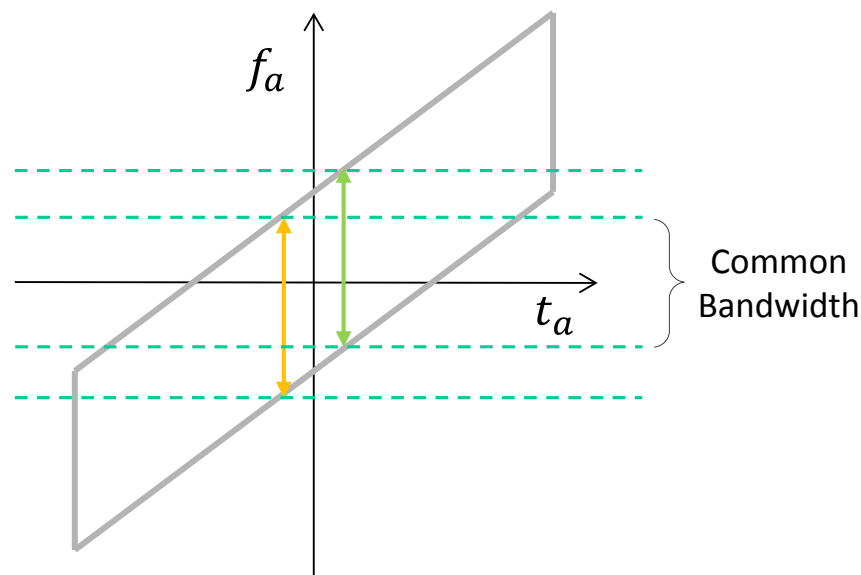
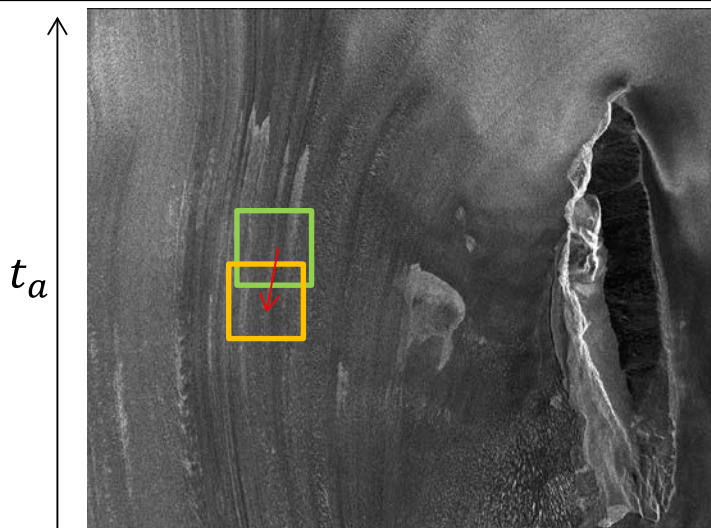
Δf_{shift} : Spectral shift

K_{rot} : Rotation rate (steering)

K_a : Target Doppler rate

v_g : Ground Velocity

Δ_{az} : Azimuth displacement (metres)



: Master Patch : Slave Patch

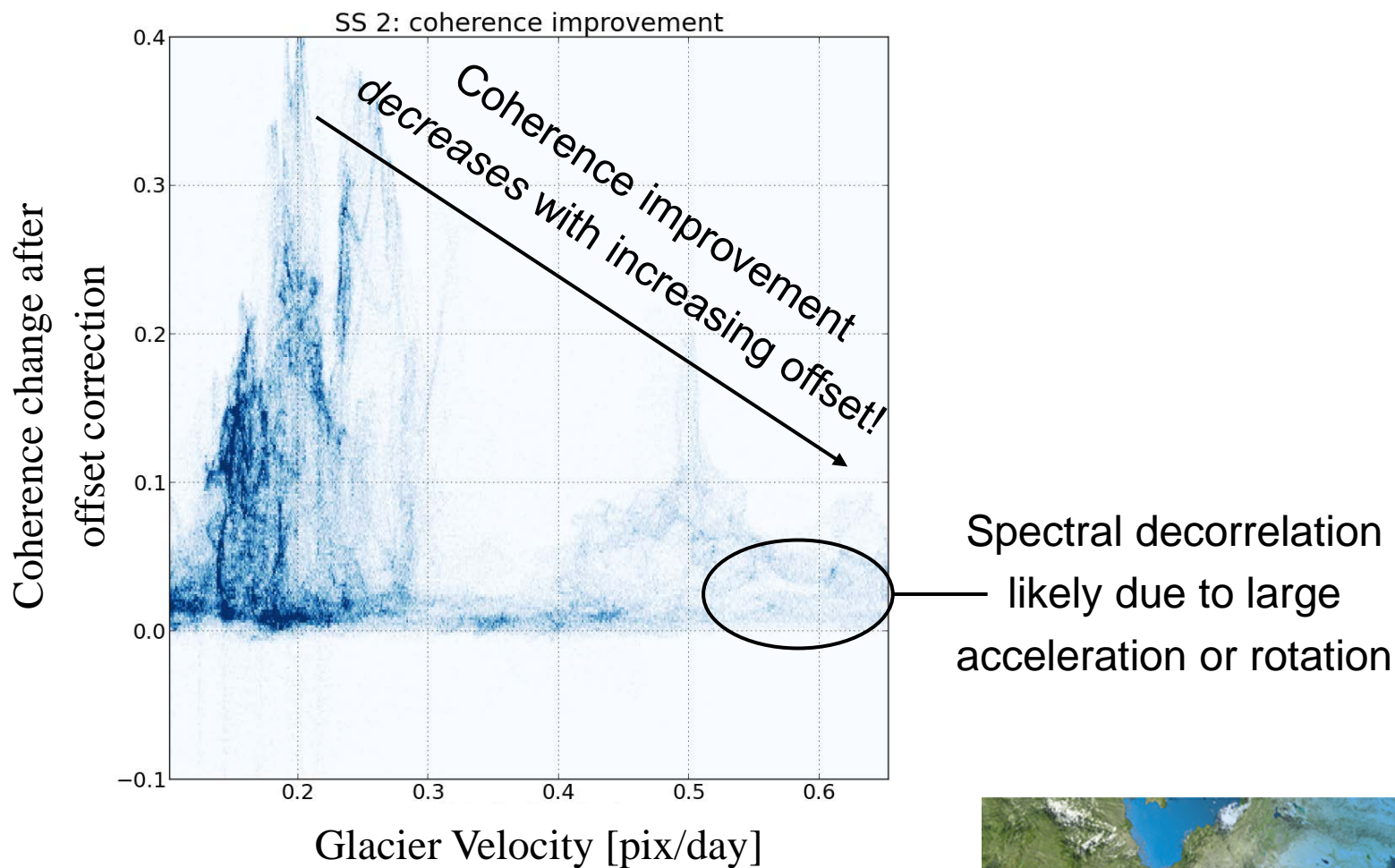
- Total decorrelation in Sentinel-1 for 1.3 km azimuth displacement (~100 az. samples).
- Three step processing: estimate + filter + re-estimate



Time-series Evaluation – Temporal decorrelation

Analysis with TerraSAR-X Data

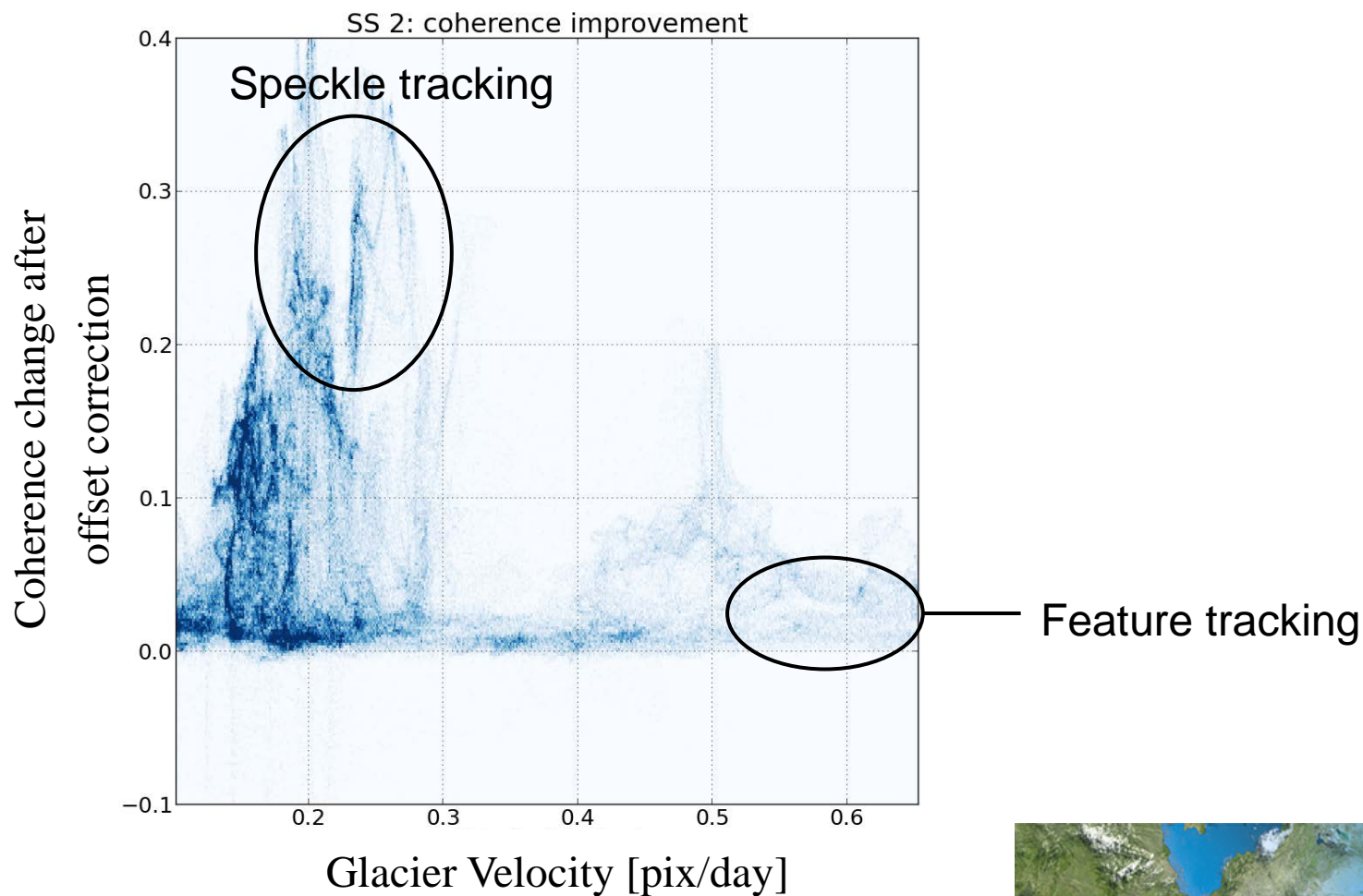
- Expectation: The larger the offset, the more coherence improves after coregistration. But...



Timeseries Evaluation – Temporal decorrelation

Analysis with TerraSAR-X Data

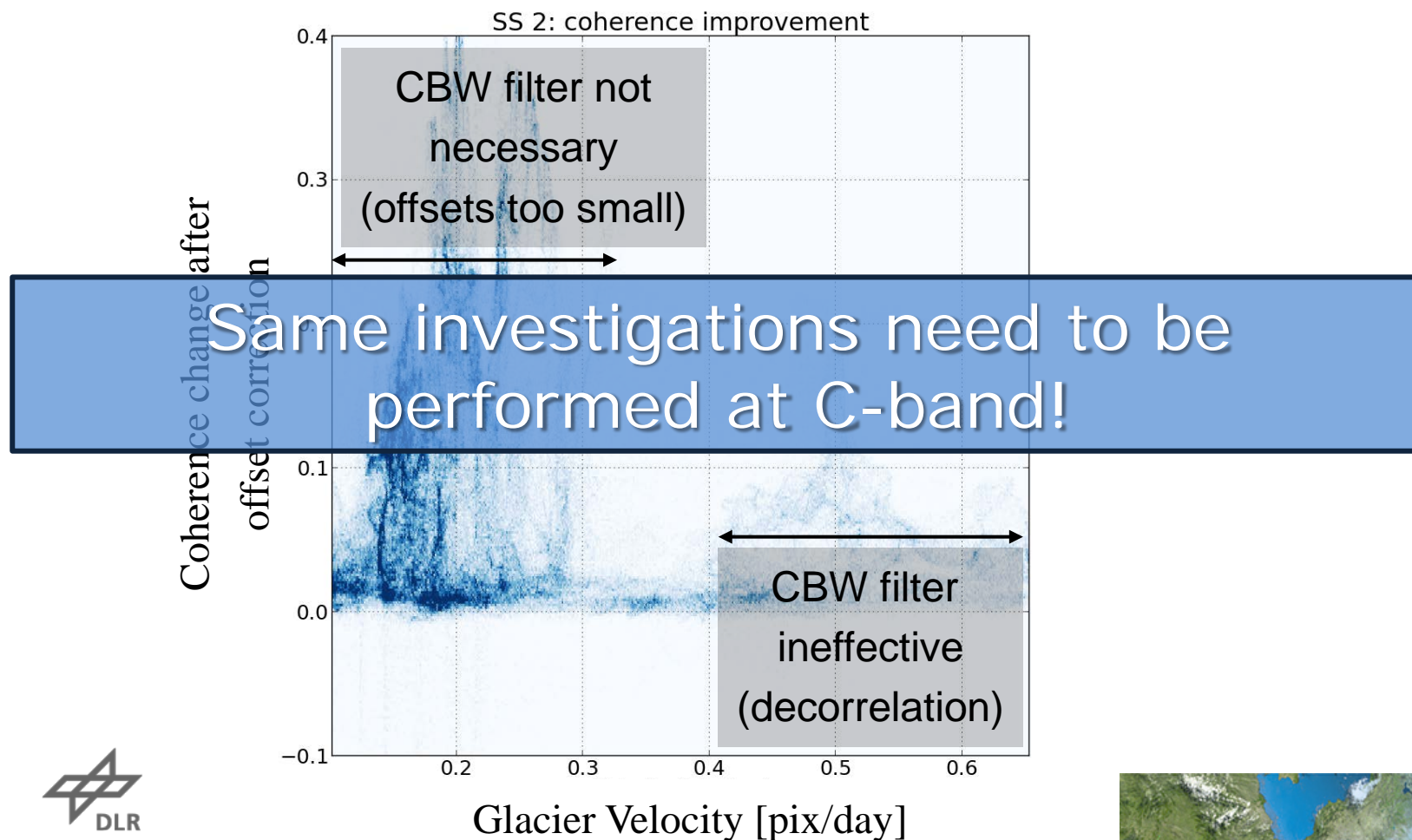
- Feature tracking occurs where the correction of large offsets yields little coherence improvement:



Timeseries Evaluation – Temporal decorrelation

Analysis with TerraSAR-X Data

- Implications for the azimuth common bandwidth filter:



Role of the Orbital Tube



The Orbital Tube

- Sentinel-1 is controlled to follow an orbital tube of 50 m radius.
- The size of the tube can have a critical role in terms of **azimuth spectral decorrelation** and **azimuth coregistration accuracy** for the TOPS mode.
- The baseline components can be described using the **Clohessy-Wiltshire** equations:

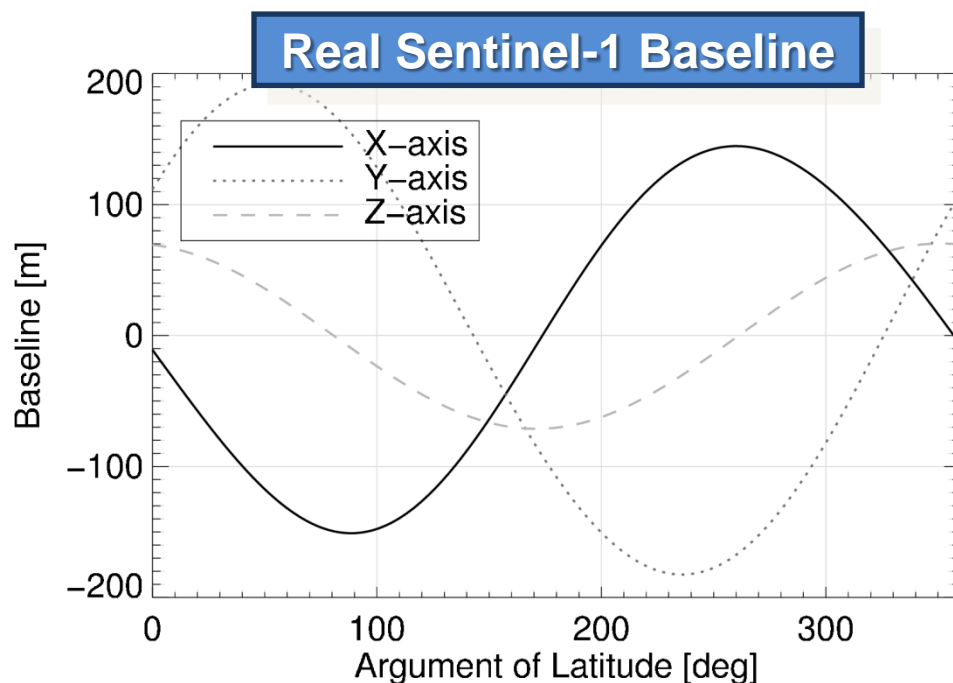
$$\Delta x(t) = \Delta x_0 + 2 \cdot A \cdot \sin(\phi_{lat} + \alpha),$$

$$\Delta y(t) = -B \cdot \cos(\phi_{lat}),$$

$$\Delta z(t) = -A \cdot \cos(\phi_{lat} + \alpha)$$

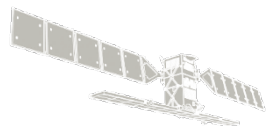
Depends on difference between orbit eccentricities

Depends on difference between inclination vectors

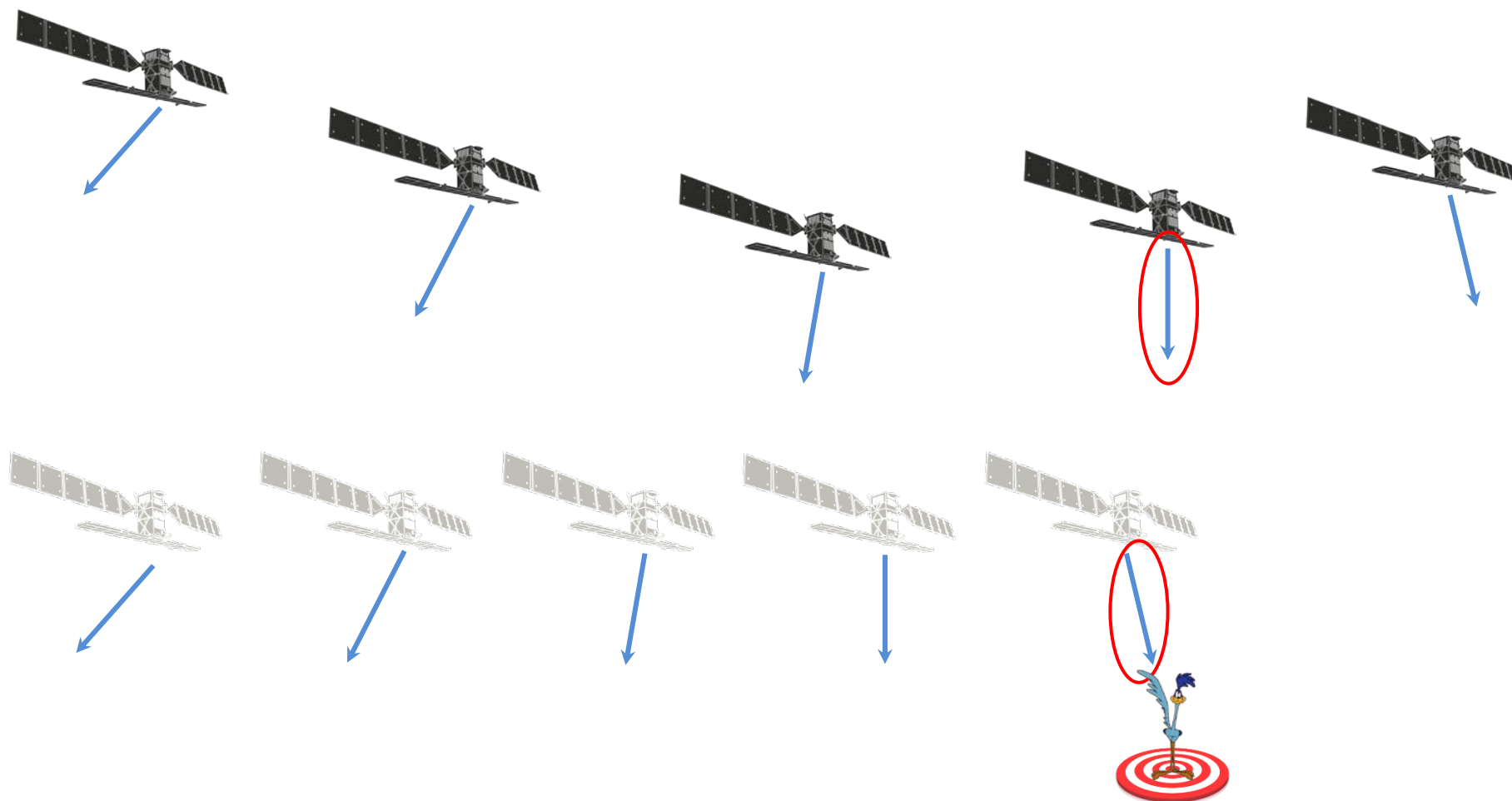


Burst Synchronization for Long Data Takes

- Sentinel-1 can synchronize with very good accuracy at the beginning of the data take (2-3 ms std. dev.).
- However, this does not guarantee perfect synchronization during the data take. Sentinel-1 can potentially acquire **25 minutes** per orbit.
- The **along-track component** of the baseline introduces azimuth spectral decorrelation, as the targets are observed under a different squint angle due to **TOPS scanning pattern**.



Burst Synchronization for Long Data Takes

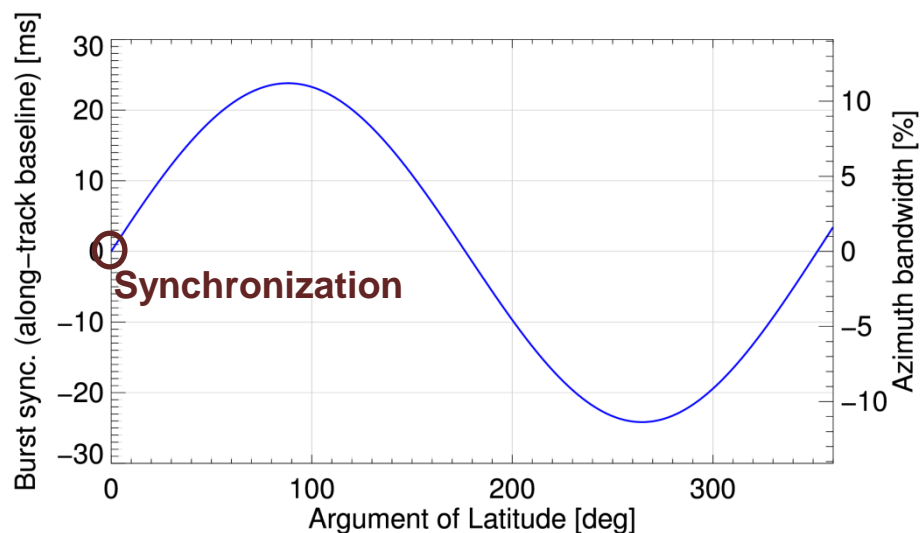


Target observed under different Doppler centroids = Azimuth spectral decorrelation!



Burst Synchronization for Long Data Takes

- Numerical evaluation using Sentinel-1 real orbits:

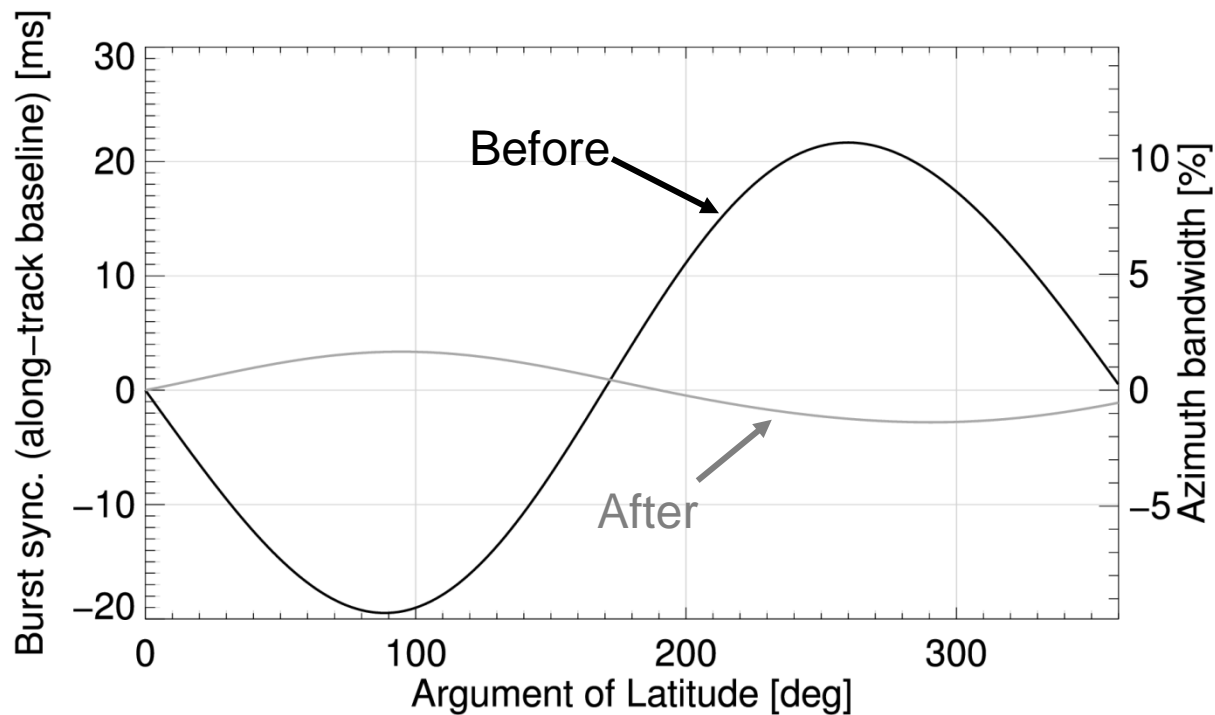


- Observations:
 - For a 25 minutes data take (a quarter of orbit), up to 30 ms mis-synchronization at the end of the data take \Rightarrow 15 % azimuth bandwidth (30% worst case in stacks!)
 - Drift depends on latitude: (± 0.4 ms/deg). Worst case at equator.
- Along-track component depends on radial tube dimension \Rightarrow Radius of radial tube, A , depends on difference between orbit eccentricities \Rightarrow Tighter orbit eccentricity control to mitigate effect (not necessarily more maneuvers or fuel!)



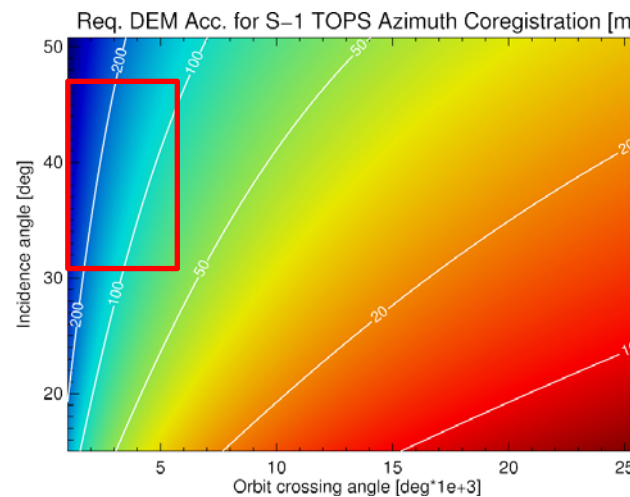
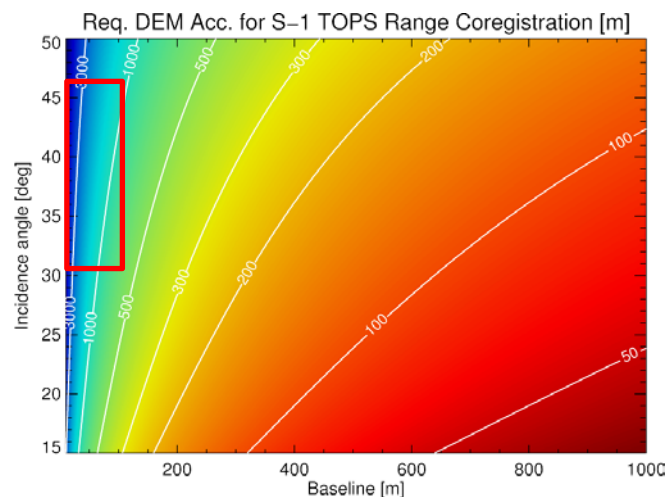
Burst Synchronization for Long Data Takes

- Reduction of radial tube size implemented in Sentinel-1:



Crossing Angle – DEM Requirements for Coregistration

- Geometric coregistration: use the external DEM to compute the offsets between image pairs (nominal coregistration), rather than estimate them from the data.
- The non-parallel orbit increases the requirement in the DEM accuracy when it is used for coregistration*.



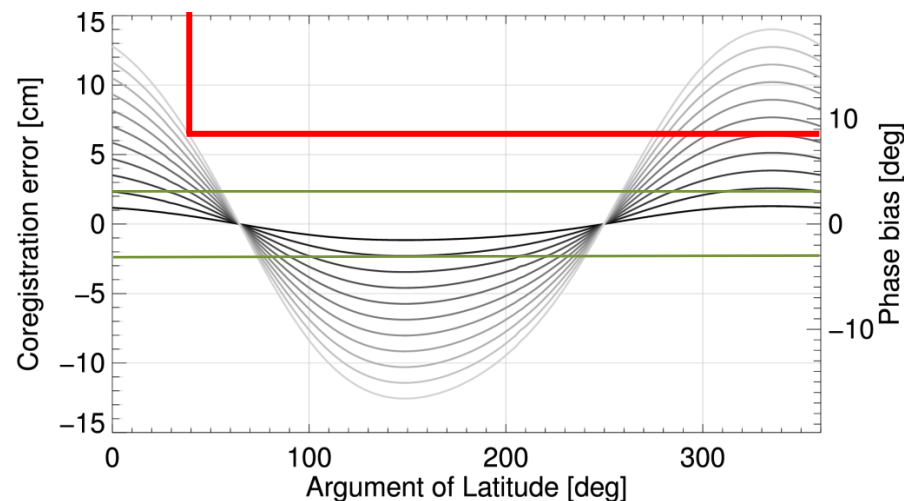
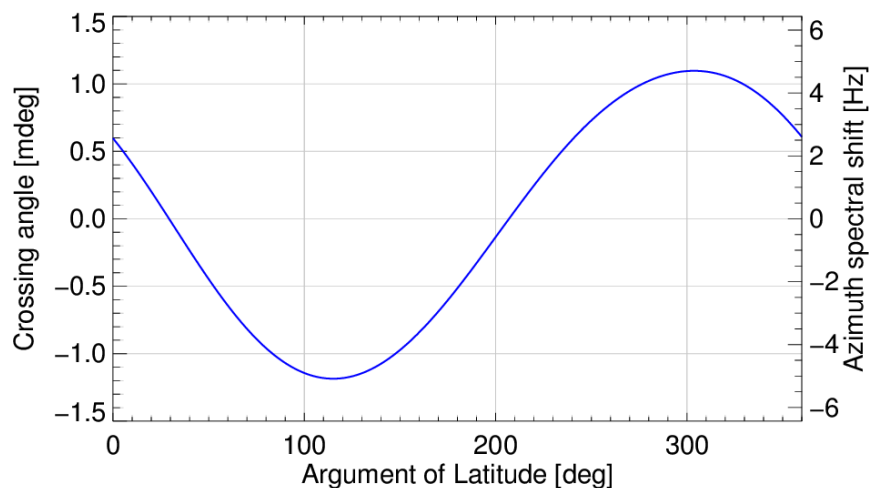
Range: Small orbital tube of S-1 (50m in diameter). No stringent requirement.

Azimuth: Small S-1 orbit crossing angles anticipated (0.001° worst case).
SRTM/ASTER DEMs sufficiently accurate!

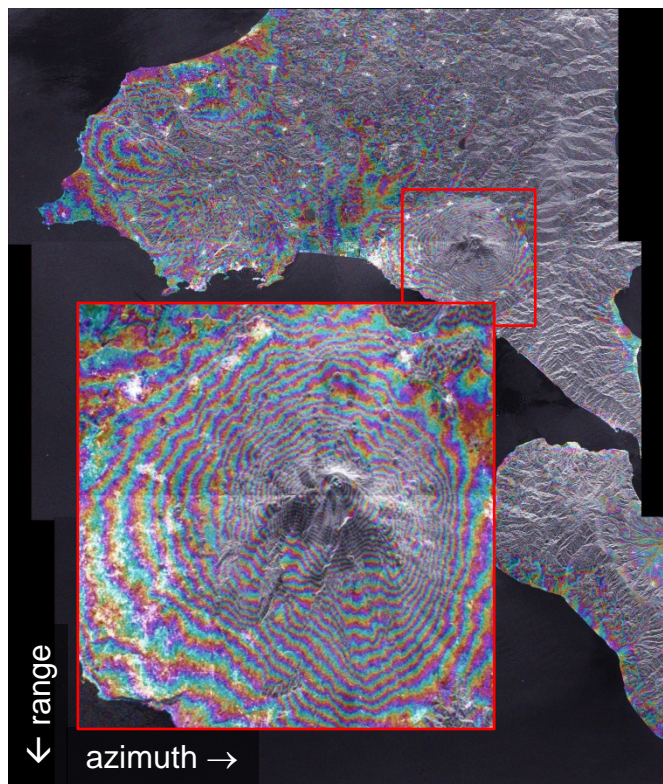
*E. Sansosti, P. Berardino, M. Manunta, F. Serafino, G. Fornaro, "Geometrical SAR Image Registration," *IEEE Trans. on Geosci. and Remote Sens.*, vol. 44, no. 10, Oct. 2006.

Crossing Angle – DEM Requirements for Coregistration

- Accurate computation of DEM requirement using Sentinel-1 orbits.
- The larger the variation in across-track (B variable in the Clohessy-Wiltshire equations), the larger the crossing angle will be.
- The plot on the right shows the azimuth coregistration error for different error heights (ranging from 300m to 3300 m) using real Sentinel-1 orbits.
- Errors in the DEM up to 600m produce less than 3° phase error. DEM requirement proportional to cross-track dimension of tube.



Case Example: Mount Etna Example



Orbit crossing angle: ~ 0.3 mdeg

Bperp (per Sub-swath): 140.5 / 129.4 / 118.2 m

Burst mis-sync: 3ms

Difference of Slaves co-registered with and without DEM (reference = 0m)

